

Scanning of stellar images by a slit diafragn
and estimation of their parameters.

I. Approximation of turbulent disks by a
Gaussian distribution

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Abstract

Exact analytical expressions are derived for evaluation of the smoothed sizes of the stellar images when they are scanned by an infinitely long slit with a finite width $2h$. Two cases are considered: (i) the data are approximated by a Gaussian curve by the least squares method; (ii) the data are approximated by a Gaussian curve, which has the same y-coordinate of the center of weight as the smoothed curve.

Key words: stellar images, scanning.

I. Introduction

It is well known that for a long enough exposure times (e.g., to smear out the speckle structure) the distribution of light intensity within the turbulent disks of stellar images may be well fitted by a Gaussian function with a dispersion σ^2 . This statement is true if the observations are made by a ground-based telescope with a large aperture, greater than ~ 50 cm, and if also a narrow transmission band filters are used to eliminate the atmospheric dispersion at large zenith angles. The conclusions following below are, obviously, applicable when photographic stellar images are scanned by a long slit window with a finite width $2h$.

Let us introduce in the focal plane of the telescope a rectangular coordinate system (x, y) with an origin at the center of the turbulent stellar disk. Then the adopted unsmoothed intensity distribution is given by the expression

$$(1) \quad g(x, y) = \frac{S}{2\pi\sigma_0^2} \exp\left(-\frac{x^2+y^2}{2\sigma_0^2}\right),$$

where S is a normalization constant and σ_0^2 is the dispersion of the two-dimensional Gaussian distribution. Adopting a circular symmetry of the stellar images we set $\sigma_x = \sigma_y = \sigma_0$. The quantity σ_0 is connected with the Fried's parameter r_0 , giving the correlation radius of the atmospheric turbulence (Fried, 1966); the greater r_0 , the smaller is σ_0 . Because we are interesting only of the relative light distribution in the stellar images and not of their total photometric fluxes we set further $S = 1$. For simplicity we set also $\sigma_0 = 1$, remembering hereafter in this paper that all linear scales are measured in unit of length σ_0 .

II. Definition of the problem

We shall describe the rectangular window of the scanning diafragn by an infinitely long slit, with a finite width $2h$. The slit is aligned along the axis y . Denoting by x the abscissa of the middle line of the slit, the intensity flux $2h\tilde{g}(x)$ transmitted by the diafragn is given as follows

$$(2) \quad \tilde{g}(x) = \int_{x-h}^{x+h} \left[\int_{-\infty}^{\infty} g(x, y) dy \right] dx = (2h)^{-1} \left[\Phi(x+h) - \Phi(x-h) \right],$$

where $\Phi(x) = (2\pi)^{-1/2} \int_{-\infty}^x \exp(-t^2/2) dt$ is the error integral (Korn et al., 1977; Janke et al., 1977) and performing the integration by y in (2) we have get the multiplier $\sqrt{2\pi}$. The smoothed distribution $\tilde{g}(x)$ is not exactly a Gaussian curve but we may approximate it by a Gaussian one with a dispersion σ^2

$$(3) \quad \tilde{g}^G(x) = (\sqrt{2\pi} \sigma)^{-1} \exp(-x^2/2\sigma^2) .$$

Of course, if the width of the slit $2h$ is not too large, we would expect that the value of σ is not very different from that of $\sigma_0 = 1$ of the unsmoothed distribution $g(x, y)$. In this paper we give exact functional relations between σ, σ_0 and h for two different cases: (i) the approximation (3) is performed by the least squares method; and (ii) the approximation (3) is obtained by equating the centres of weight (by the y -coordinates) of the areas constrained by the unsmoothed and smoothed curves, respectively. These results may be used to correct the measured σ 's for the finite width of the slit $2h$, when very precise values of σ_0 's are needed to be known. Obviously, the results in this paper are applicable not only in the case of a real-time astronomical scanning of stellar images. They are useful in the slit microphotometry of photographic stellar images.

III. Approximation of $\tilde{g}(x)$ with a Gaussian curve by the least squares method

In this section we shall denote \tilde{g}^G from (3) by \tilde{g}_{LS} .

The approximation by the method of least squares in our case is described by the condition (Brandt, 1975)

$$(4) \int_{-\infty}^{\infty} \left[\tilde{g}(x) - \tilde{g}_{LS}(x) \right]^2 dx = \min ,$$

or equivalently

$$(5) \int_{-\infty}^{\infty} \left[\tilde{g}(x) - \tilde{g}_{LS}(x) \right] \frac{d}{d\sigma} \tilde{g}_{LS}(x) dx = 0.$$

This is an equation for the desired quantity σ as a function of $2h$ and σ_0 (in units $\sigma_0 = 1$). Substituting (2) and (3) in (5), the least squares condition for σ becomes

$$(6) \int_{-\infty}^{\infty} \left[\Phi(x+h) - \Phi(x-h) - \frac{2h}{\sqrt{2\pi}\sigma} \exp\left(-\frac{x^2}{2\sigma^2}\right) \right] \\ \times (x^2 - \sigma^2) \exp\left(-\frac{x^2}{2\sigma^2}\right) dx = 0.$$

Further we denote by $\mathcal{K}_0(h)$ the integral

$$(7) \mathcal{K}_0(h) = \int_{-\infty}^{\infty} \Phi(x+h) \exp(-x^2/2\sigma^2) dx ,$$

and rewrite (6) as a sum of six integrals

$$(8) \sum_{i=1}^6 \mathcal{K}_i(h) = 0 ,$$

where $\mathcal{K}_i(h)$ ($i = 1, \dots, 6$) are defined further in this section as follows

$$(9) \mathcal{K}_1(h) = \int_{-\infty}^{\infty} \Phi(x+h) x^2 \exp(-x^2/2\sigma^2) dx .$$

Integrating by parts we obtain that

$$(10) \mathcal{K}_1(h) = \sigma^2 \mathcal{K}_0(h) + \frac{\sigma^2}{\sqrt{2\pi}} \left\{ \int_{-\infty}^{\infty} x \exp \left[-\left(\frac{1+\sigma^2}{2\sigma^2}\right)x^2 + 2\left(-\frac{h}{2}\right)x \right] dx \right\} \exp\left(-\frac{h^2}{2}\right).$$

Gradshteyn et al.(1963) (formula 3.462.2) have evaluated the integral

$$(11) \int_{-\infty}^{\infty} x^n \exp(-px^2 + 2qx) dx \\ = \frac{1}{2^{n-1} p} \sqrt{\frac{\pi}{p}} \frac{d^{n-1}}{dq^{n-1}} \left[q \exp\left(\frac{q^2}{p}\right) \right]$$

for n positive integer and $p > 0$. In our case for $n = 1$, $p = (1 + \sigma^2)/2\sigma^2$ and $q = -h/2$ we get finally, according to (10)

$$(12) \mathcal{K}_1(h) = \sigma^2 \mathcal{K}_0(h) - \frac{\sigma^5 h}{(1 + \sigma^2)^{3/2}} \exp \left[-\frac{h^2}{2(1 + \sigma^2)} \right],$$

$$(13) \mathcal{K}_2(h) = -\sigma^2 \int_{-\infty}^{\infty} \Phi(x+h) \exp(-x^2/2\sigma^2) dx = -\sigma^2 \mathcal{K}_0(h),$$

$$(14) \mathcal{K}_3(h) = -\int_{-\infty}^{\infty} \Phi(x-h) x^2 \exp(-x^2/2\sigma^2) dx = -\mathcal{K}_1(-h),$$

$$(15) \mathcal{K}_4(h) = \sigma^2 \int_{-\infty}^{\infty} \Phi(x-h) \exp(-x^2/2\sigma^2) dx = \sigma^2 \mathcal{K}_0(-h),$$

and

$$(16) \quad \mathcal{K}_5(h) = -\frac{2h}{\sqrt{2\pi}} \frac{1}{\sigma} \int_{-\infty}^{\infty} x^2 \exp(-x^2/\sigma^2) dx.$$

Integrating by parts and using the result that the error integral at infinity is equal to unity (e.g., $\Phi(\infty) = 1$), we obtain

$$(17) \quad \mathcal{K}_5(h) = -h\sigma^2/\sqrt{2}.$$

Finally we get

$$(18) \quad \mathcal{K}_6(h) = \frac{2h\sigma}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \exp(-x^2/\sigma^2) dx = 2h\sigma^2/\sqrt{2}.$$

Substituting the expressions (12) - (15), (17) and (18) into the least squares condition (8), and performing the summation we obtain the desired functional dependence between the slit width $2h$ and the dispersion σ^2 of the Gaussian distribution by which the smoothed curve $\tilde{g}(x)$ from (2) has been approximated. Note that the integrals $\mathcal{K}_6(h)$ and $\mathcal{K}_6(-h)$ during summing up annihilate, so there is not need to compute them.

$$(19) \quad \frac{\sigma^3}{(1+\sigma^2)^{3/2}} \exp\left[-\frac{h^2}{2(1+\sigma^2)}\right] = \frac{1}{2\sqrt{2}}.$$

This relation uniquely determines the dependence between h and σ because the multipliers $\sigma^3/(1+\sigma^2)^{3/2}$ and $\exp[-h^2/2(1+\sigma^2)]$ for every one fixed value of h are strict monotonically increasing functions of σ , and also the exponential factor strictly decreases with increasing of h

(for a fixed value of σ). It is more useful to express h as a function of σ . For this purpose we may take the natural logarithms from two sides of (19) and the desired final expression is

$$(20) \quad 2h = 2 \left[(1 + \sigma^2) \ln \frac{8\sigma^6}{(1 + \sigma^2)^3} \right]^{1/2} .$$

Further the quantity σ , connected with h through relations (19) or (20) (obtained by the least squares method), is denoted by σ_{LS} . In Figures 1 and 2 the dependence of $[\sigma_{LS}(h) - \sigma_0] / \sigma_0$ vs. $2h/\sigma_0$ is plotted by solid lines. As mentioned earlier, in the above calculations σ_0 has been set equal to 1 .

IV. Approximation of $\tilde{g}(x)$ with a Gaussian curve by the y-coordinate of the center of weight

It is easy to check that the one-dimensional Gaussian curve $\tilde{g}^G(x)$ (3) has an ordinate y_G of the center of weight of the area \int ($\int = 1$ in our case) given by (Fihtengoltz, 1969)

$$(21) \quad y_G = \frac{\int}{4\sqrt{\pi} \sigma} ; \quad (\int = 1) .$$

We shall now compute the same quantity y_G , but for the smoothed curve $\tilde{g}(x)$ given by (2), which is not exactly Gaussian one, and by equating this value to y_G from (21) we shall obtain an estimation for σ . The smoothed "observed" curve $\tilde{g}(x)$ is supposed to be symmetric relatively the axis $x = 0$, and computing y_G we shall restrict us only to the case $x \geq 0$. The y-coordinate of the center of weight of the smoothed curve

$\tilde{g}(x)$ may be expressed as a ratio of two functions of h

$$(22) \quad y_G = A_1(h) / A_2(h) ,$$

where

$$(23) \quad A_1(h) = \int_0^{\tilde{g}^{(0)}} x(\tilde{g}) \tilde{g} d\tilde{g} ,$$

and

$$(24) \quad A_2(h) = \int_0^{\tilde{g}^{(0)}} x(\tilde{g}) d\tilde{g} .$$

From (2) we also have

$$(25) \quad d\tilde{g} = \frac{1}{\sqrt{2\pi} 2h} \left\{ \exp\left[-\frac{(x+h)^2}{2}\right] - \exp\left[-\frac{(x-h)^2}{2}\right] \right\} dx .$$

As follows from (2), (23) and (25)

$$(26) \quad A_1(h) = (2\pi)^{-1/2} (2h)^{-2} \int_0^{\infty} x \left[\Phi(x+h) - \Phi(x-h) \right] \\ \times \left\{ \exp\left[-\frac{(x-h)^2}{2}\right] - \exp\left[-\frac{(x+h)^2}{2}\right] \right\} dx .$$

Let us denote by $L_0(h)$ the expression

$$(27) \quad L_0(h) = \int_{-\infty}^{\infty} \Phi(x+h) \exp(-x^2/2) dx .$$

In the following we shall also use the property of the error integral $\Phi(x)$ (Janke et al., 1977)

$$(28) \quad \Phi(x) = 1 - \Phi(-x) .$$

Obviously, according to (26), the quantity $(2\pi)^{1/2} (2h)^2 A_1(h)$ may be expressed as a sum of the four integrals $L_i(h)$ ($i = 1, \dots, 4$),

defined by the next relations

$$\begin{aligned}
 (29) \quad L_1(h) &= \int_0^{\infty} x \Phi(x+h) \exp \left[-(x-h)^2/2 \right] dx \\
 &= - \int_0^{\infty} \Phi(x+h) d \exp \left[-(x-h)^2/2 \right] \\
 &\quad + h \int_0^{\infty} \Phi(x+h) \exp \left[-(x-h)^2/2 \right] dx .
 \end{aligned}$$

Integrating by parts the first term in the right-hand side of the above expression, finally we obtain

$$\begin{aligned}
 (30) \quad L_1(h) &= \Phi(h) \exp(-h^2/2) + (2\sqrt{2})^{-1} \exp(-h^2/2) \\
 &\quad + h \int_0^{\infty} \Phi(x+h) \exp \left[-(x-h)^2/2 \right] dx .
 \end{aligned}$$

Further we have

$$\begin{aligned}
 (31) \quad L_2(h) &= - \int_0^{\infty} x \Phi(x+h) \exp \left[-(x+h)^2/2 \right] dx \\
 &= \int_0^{\infty} \Phi(x+h) d \exp \left[-(x+h)^2/2 \right] \\
 &\quad + h \int_0^{\infty} \Phi(x+h) \exp \left[-(x+h)^2/2 \right] dx .
 \end{aligned}$$

By an analogy with $L_1(h)$, the final result is

$$(32) \quad L_2(h) = - \Phi(h) \exp(-h^2/2) - (2\pi)^{-1/2} \int_0^{\infty} \exp(x+h)^2 dx$$

$$+h \int_0^{\infty} \Phi(x+h) \exp\left[-(x+h)^2/2\right] dx,$$

$$(33) L_3(h) = - \int_0^{\infty} x \Phi(x-h) \exp\left[-(x-h)^2/2\right] dx = L_2(-h),$$

and

$$(34) L_4(h) = \int_0^{\infty} x \Phi(x-h) \exp\left[-(x+h)^2/2\right] dx = L_1(-h).$$

Consequently

$$(35) \sum_{i=1}^4 L_i(h) = (2)^{-1/2} \exp(-h^2) + h \left[M_1(h) + M_2(h) \right] - (2\pi)^{-1/2} \left\{ \int_0^{\infty} \exp\left[-(x+h)^2\right] dx + \int_0^{\infty} \exp\left[-(x-h)^2\right] dx \right\},$$

where the notations

$$(36) M_1(h) = \int_0^{\infty} \Phi(x+h) \exp\left[-\frac{(x-h)^2}{2}\right] dx - \int_0^{\infty} \Phi(x-h) \exp\left[-\frac{(x+h)^2}{2}\right] dx,$$

and

$$(37) M_2(h) = \int_0^{\infty} \Phi(x+h) \exp\left[-\frac{(x+h)^2}{2}\right] dx - \int_0^{\infty} \Phi(x-h) \exp\left[-\frac{(x-h)^2}{2}\right] dx$$

are used. Using the property (28) it may be shown that

$$(38) \int_0^{\infty} \Phi(x-h) \exp\left[-\frac{(x+h)^2}{2}\right] dx = \int_0^{\infty} \left[1 - \Phi(h-x)\right] \exp\left[-\frac{(x+h)^2}{2}\right] dx \\ = \int_{-\infty}^{-h} \exp(-x^2/2) dx - \int_{-\infty}^0 \Phi(x+h) \exp\left[-(x-h)^2/2\right] dx.$$

It is easy to see that

$$(39) M_1(h) = \int_{-\infty}^{\infty} \Phi(x+2h) \exp(-x^2/2) dx - \sqrt{2\pi} \Phi(-h) \\ = L_0(2h) - \sqrt{2\pi} \Phi(-h).$$

Analogically, using again the property (28), the second term in the square brackets in the right-hand side of (35) may be computed as follows

$$(40) \quad M_2(k) = \int_{-\infty}^{\infty} \Phi(x) \exp(-x^2/2) dx - (2\pi)^{1/2} \left[1 - \Phi(-k) \right] \\ = L_0(0) - \sqrt{2\pi} + \sqrt{2\pi} \Phi(-k).$$

Substituting the results (39) and (40) into (35), and performing the transition from the variable x to $x\sqrt{2}$, and denoting also by $k' = k\sqrt{2}$ in the brace parentheses of the last term of (35), we find that

$$(41) \quad (2\pi)^{1/2} (2k)^2 A_1(k) \equiv \sum_{i=1}^4 L_i(k) \\ = (2)^{-1/2} \exp(-k^2) + k \left[L_0(2k) + L_0(0) - \sqrt{2\pi} \right] \\ - (4\pi)^{-1/2} \left[\mathcal{N}_1(k') + \mathcal{N}_2(k') \right],$$

where

$$(42) \quad \mathcal{N}_1(k) = \int_0^{\infty} \exp \left[-(x+k)^2/2 \right] dx,$$

and

$$(43) \quad \mathcal{N}_2(k) = \int_0^{\infty} \exp \left[-(x-k)^2/2 \right] dx = \mathcal{N}_1(-k).$$

There is not difficulties to find analytical expressions for the integrals $\mathcal{N}_1(k')$ and $\mathcal{N}_2(k')$.

$$(44) \quad \mathcal{N}_1(k') = \sqrt{2\pi} \Phi(-k').$$

According to (28)

$$(45) \quad \mathcal{N}_1(k') = \sqrt{2\pi} - \sqrt{2\pi} \Phi(k') .$$

Therefore

$$(46) \quad (4\pi)^{-1/2} \left[\mathcal{N}_1(k') + \mathcal{N}_2(k') \right] = (2)^{-1/2} .$$

Returning to the definition (27) and differentiating this expression with respect to k , we have

$$(47) \quad \frac{dL_0(k)}{dk} = \frac{\exp(-k^2/2)}{\sqrt{2\pi}} \int_{-\infty}^{\infty} \exp(-x^2 - kx) dx .$$

According to Gradshteyn et al. (1963) (formula 3.323.2)

$$(48) \quad \int_{-\infty}^{\infty} \exp(-p_1^2 x^2 \pm q_1 x) dx = \frac{\sqrt{\pi}}{p_1} \exp\left(\frac{q_1^2}{4p_1^2}\right)$$

for $p_1 > 0$.

Substituting in the above expression $p_1 = 1$ and $q_1 = k$, we obtain the following differential equation for $L_0(k)$

$$(49) \quad \frac{dL_0(k)}{dk} = \frac{1}{\sqrt{2}} \exp(-k^2/4) .$$

Its solution is equal to

$$(50) \quad L_0(k) = \sqrt{2\pi} \Phi(k/\sqrt{2}) + C ,$$

where C is an arbitrary constant. Using the condition that the error integral is equal to unity at infinity (e.g., $\Phi(\infty) = 1$), from (50) follows that $L_0(\infty) = \sqrt{2\pi} + C$. On the other hand from (27) we have $L_0(\infty) = \sqrt{2\pi}$. Consequently, $C = 0$. Taking into account that $\Phi(0) = 1/2$, we may write

$$(51) \quad L_0(0) = \sqrt{\pi/2} \quad ,$$

$$(52) \quad L_0(2k) = \sqrt{2\pi} \Phi(k\sqrt{2}) \quad ,$$

and

$$(53) \quad L_0(2k) + L_0(0) - \sqrt{2\pi} = \sqrt{2\pi} \Phi(k\sqrt{2}) - \sqrt{\pi/2} \quad .$$

Substituting (46) and (53) into (41) we obtain the final expression for $A_1(k)$

$$(54) \quad A_1(k) = (2\pi)^{-1/2} (2k)^{-2} \left\{ \frac{1}{\sqrt{2}} \exp(-k^2) - \frac{1}{\sqrt{2}} + k \left[\sqrt{2\pi} \Phi(k\sqrt{2}) - \sqrt{\pi/2} \right] \right\} \quad .$$

Using (25), the expression (24) for $A_2(k)$ may be rewritten as

$$(55) \quad A_2(k) = -(2k\sqrt{2\pi})^{-1} \left[P_1(k) - P_2(k) \right] \quad ,$$

where

$$(56) \quad P_1(k) = \int_0^{\infty} x \exp \left[-(x+k)^2/2 \right] dx \quad ,$$

and

$$(57) \quad P_2(k) = \int_0^{\infty} x \exp \left[-(x-k)^2/2 \right] dx \quad .$$

It is easy to check that

$$(58) \quad P_1(k) = \exp(-k^2/2) - k N_1(k) \quad ,$$

and

$$(59) \quad P_2(k) = P_1(-k) = \exp(-k^2/2) + k N_1(-k) \quad .$$

Consequently, taking into account (42), (43) and (28)

$$(60) \quad A_2(h) = (2h\sqrt{2\pi})^{-1} h \left[\mathcal{N}_1(h) + \mathcal{N}_1(-h) \right] = 1/2.$$

Of course, this result would be expected from our assumption $\mathcal{S} = 1$, as mentioned earlier, under the circumstance that the integral (24) gives the half of the area enclosed within $\tilde{g}(x)$ and the abscissa ($y = 0$), and the ordinate ($x = 0$). Equating the right-hand sides of the expressions (21) and (22) we can obtain the value of σ , which we shall denote by σ_{yG} . This means that we have approximated the smoothed quasi-gaussian curve $\tilde{g}(x)$ with an exactly Gaussian curve which has the same y-coordinate of the center of weight as the $\tilde{g}(x)$ do (and also $\mathcal{S} = 1$). The final expression for σ_{yG} , according to (55) and (60), is

$$(61) \quad \sigma_{yG} = \frac{h^2}{\sqrt{2}} \left\{ \frac{\exp(-h^2) - 1}{\sqrt{2}} + h \left[\sqrt{2\pi} \Phi(h\sqrt{2}) - \sqrt{\frac{\pi}{2}} \right] \right\}^{-1}.$$

One may check that the passage to the limit $h = 0$ in the above expression gives a reasonable result, e.g., $\lim_{h \rightarrow 0} \sigma_{yG}(h) = 1$ (in view of the substitution $\sigma_0 = 1$). This check may be performed by using the L'Hospital's rule for evaluation of indeterminate expressions in terms of the ratio 0/0. In Figures 1 and 2 the dependence of $[\sigma_{yG}(h) - \sigma_0] / \sigma_0$ vs. $2h/\sigma_0$ is plotted by dashed lines. At first glance, the analytical expressions (19) (or (20)) and (61) giving the dependences between the smoothed values of σ and $2h$ for two different methods of approximation of $\tilde{g}(x)$ are very different. But as can be seen from Figures 1 and 2, the solid and the dashed lines

are very close each to other. This result is not surprising because if the width of the scanning slit $2h$ is not too large, the width of the smoothed curve σ is not too different from the width of the unsmoothed curve σ_0 . This is not a new conclusion, but the aim of this paper is to get quantitative expressions giving exact analytical connections between σ and σ_0 depending on $2h$. In real-time astronomical observations the scanning of turbulent stellar images by a slit "enlarges" their sizes typically by an amount of 0.5 - 1.5 per cent. Such an accuracy of the estimation of σ_0 is useful if a reconstruction of long-exposure turbulent astronomical images is applied.

V. Discussion

The formula (20) is obtained without the assumption that σ is a priori close to σ_0 . In the later case this expression may be simplified. Taking into account that $\sigma \gtrsim 1$ ($\sigma_0 = 1$) and using the expansion of the logarithmic function in a series we may write the following approximate expansion (Korn et al., 1977)

$$(62) \quad \ln \frac{2\sigma^2}{1+\sigma^2} \approx -1 + \frac{2\sigma^2}{1+\sigma^2} = \frac{\sigma^2 - 1}{\sigma^2 + 1},$$

and substituting this result into (20) to obtain

$$(63) \quad h = \left[3(1+\sigma_{LS}^2) \ln \frac{2\sigma_{LS}^2}{1+\sigma_{LS}^2} \right]^{1/2} \approx \sqrt{3(\sigma_{LS}^2 - 1)}.$$

Remembering that in the above formula σ_{LS} and h are measured in units of σ_0 (e.g., it has been assumed that $\sigma_0 = 1$),

(63) may be rewritten as

$$(64) \quad \sigma_{L\mathcal{S}} = (\sigma_0^2 + h^2/3)^{1/2},$$

where σ_0 , $\sigma_{L\mathcal{S}}$ and h are evaluated in ordinary units (angular or linear).

Instead of the y_G -coordinate of the area \mathcal{S} enclosed by the Gaussian curve $\tilde{g}^G(x)$ it is also possible to use the x -coordinate x_G of the center of weight of this area for $x \geq 0$ (or $|x_G|$ for $x \leq 0$). In place of formula (21) in this case we are able to introduce a new relation which includes the unknown parameter σ (Fihngoltz, 1969)

$$(65) \quad x_G = \sqrt{2/\pi} \sigma.$$

Note that in the above expression the area \mathcal{S} does not appear, e.g., this evaluation is independent of the normalization of $\tilde{g}^G(x)$. Nevertheless, when this coordinate x_G must be computed from the smoothed "observed" curve $\tilde{g}(x)$ it is necessary to decompose the area \mathcal{S} into two areas with equal sizes $\mathcal{S}_1(x \geq 0) = \mathcal{S}_2(x \leq 0) = \mathcal{S}/2$. Then we have to estimate the $|x|$ -coordinate of the center of weight for every one part of the area \mathcal{S} and to take the arithmetic mean value. So the estimation of x_G will be to some extent dependent from the way by which the area \mathcal{S} (enclosed within the smoothed curve $\tilde{g}(x)$ and the axis x) is separated. Such a procedure of decomposition is absent when we are dealing with the determination of σ through the y_G -coordinate. For this reason in this paper we have preferred to evaluate σ (by the method of the center of weight) using only the formula (21) but not the expression (65).

The author would like to thank Drs. Yu.V.Alexandrov and V.N.Dudinov from the K'harkov State University, Ukraina, for the stimulating and helpful discussions and comments.

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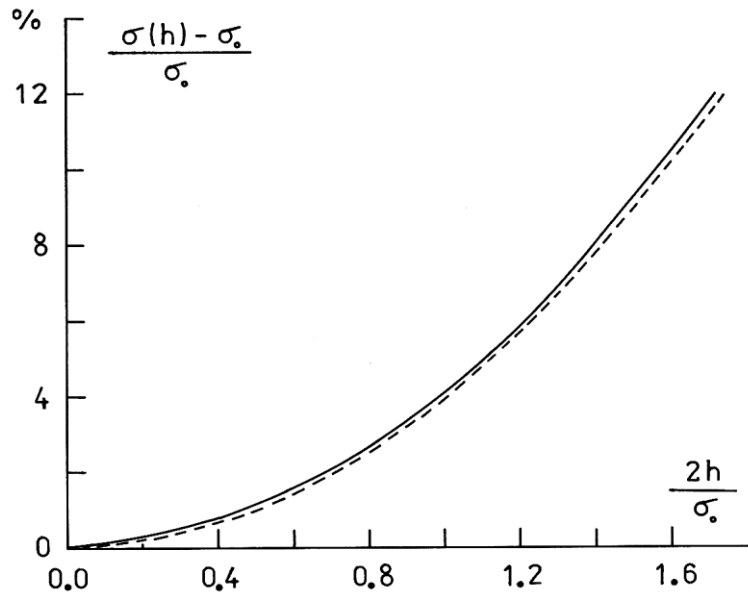


Fig.1. The dependence of $[\sigma_{LS}(h) - \sigma_0] / \sigma_0$ vs. $2h / \sigma_0$ (solid line) and $[\sigma_{yG}(h) - \sigma_0] / \sigma_0$ vs. $2h / \sigma_0$ (dashed line)

Fig.2. Same as Fig.1, but the scale is enlarged

