

Rubin's LSST as a large-scale time-domain photometric survey

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Stellar variability, stellar multiplicity: periodicity in time & motion , Sofia, June 8, 2023

Science motivation for undertaking the Legacy Survey of Space and Time

More details about science drivers and system design:
Ivezic et al. (2019):
ls.st/lop

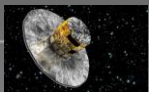
Expansion and history of the Universe and the growth of structure
(dark matter, dark energy, cosmology, spatial distribution of galaxies, gravitational lensing, supernovae): “Was Einstein right?”

Time domain: what changes on the sky?
(cosmic explosions, variable stars, unknown unknowns)

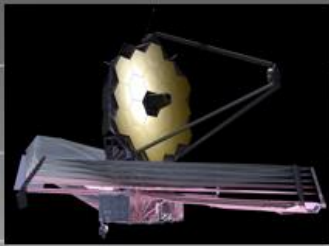
The Solar System structure
(near-Earth hazardous asteroids, main-belt asteroids, trans-Neptunian objects, comets)

The Milky Way structure
(stars as tracers of the structure and evolution of our Galaxy, interstellar matter, the physics of stars)

A key point: most of science programs will utilize the same dataset.



Gaia



James Webb
Space Telescope
(6.5m)



Nancy Grace
Roman Space
Telescope (2.4m)

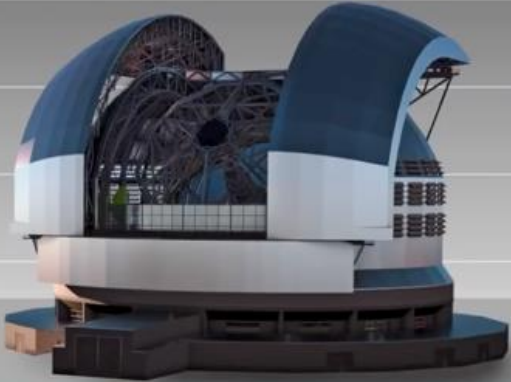
100 m

80 m

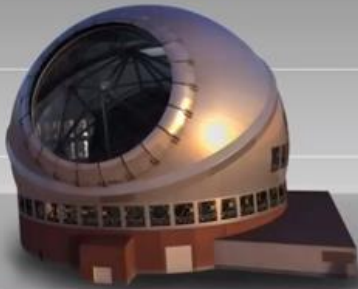
60 m

40 m

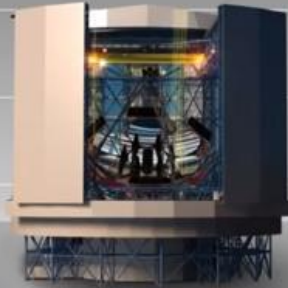
20 m



ELT: ~40m



TMT: ~30m



GMT: ~25m



Rubin: 8.4m

Rubin Obs. will not have the largest mirror but will have by far the largest product of the mirror area and the field-of-view size (etendue or throughput)



Vera C. Rubin (1928-2016)

Basic idea behind LSST: a uniform sky survey

- 90% of time will be spent on a uniform survey: every 3-4 nights, the whole observable sky will be scanned twice per night (ugrizy bands, 0.3-1.1 micron)
- in 10 years, half of the sky will be imaged about 1000 times: a digital color movie of the sky, 5-sigma r mag ~ 24
- ~100 petabytes, or 100,000 terabytes, of data: about a billion 16 megapixel images, enabling measurements for 40 billion objects, **coadded map depth 5-sigma r mag ~ 27**

SDSS

gri

3.5'x3.5'

r~22.5

3 arcmin
is 1/10 of
the full
Moon's
diameter

LSST's
field of
view is
3000
times
larger



HSC
gri
3.5'x3.5'
r~27

3 arcmin
is 1/10 of
the full
Moon's
diameter

like
LSST
depth
(but tiny
area)

LSST
will
deliver 5
million
such
images

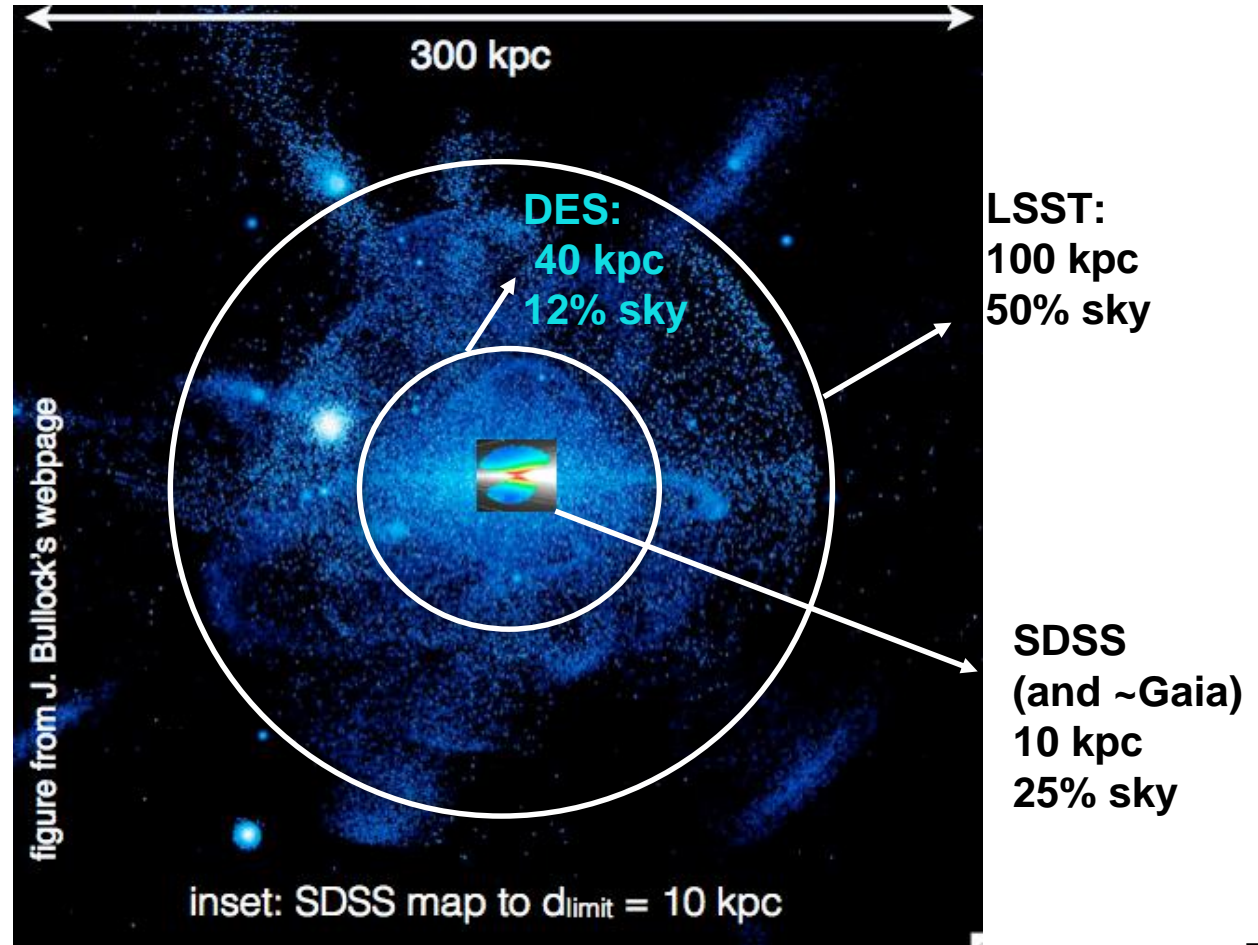


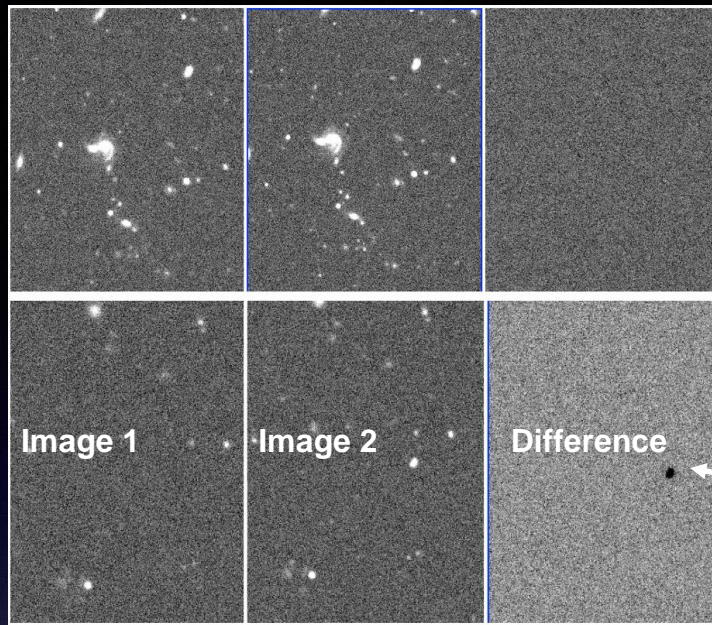
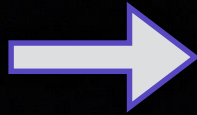
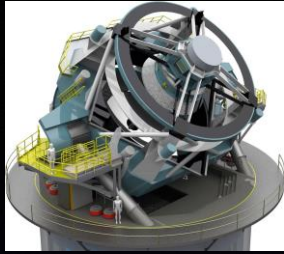
Milky Way science with upcoming coadded LSST data

Distance limits for
(very numerous!)
turn-off stars:

A few hundred million turn-off stars (with $[Fe/H]!$) out to ~ 100 kpc.

We will go beyond the limit of 100 kpc for main sequence stars with luminous variable stars RR Lyrae!





Transients with LSST

Alert!

Additional “followup” data obtained to:

- confirmation and classification
- provide better temporal resolution
- use different filters/wavelengths
- obtain spectra (distance!)
- other measurements (e.g. polarimetry)

Image 1

Image 2

Difference

Alerts can trigger “Follow-up” observations: 

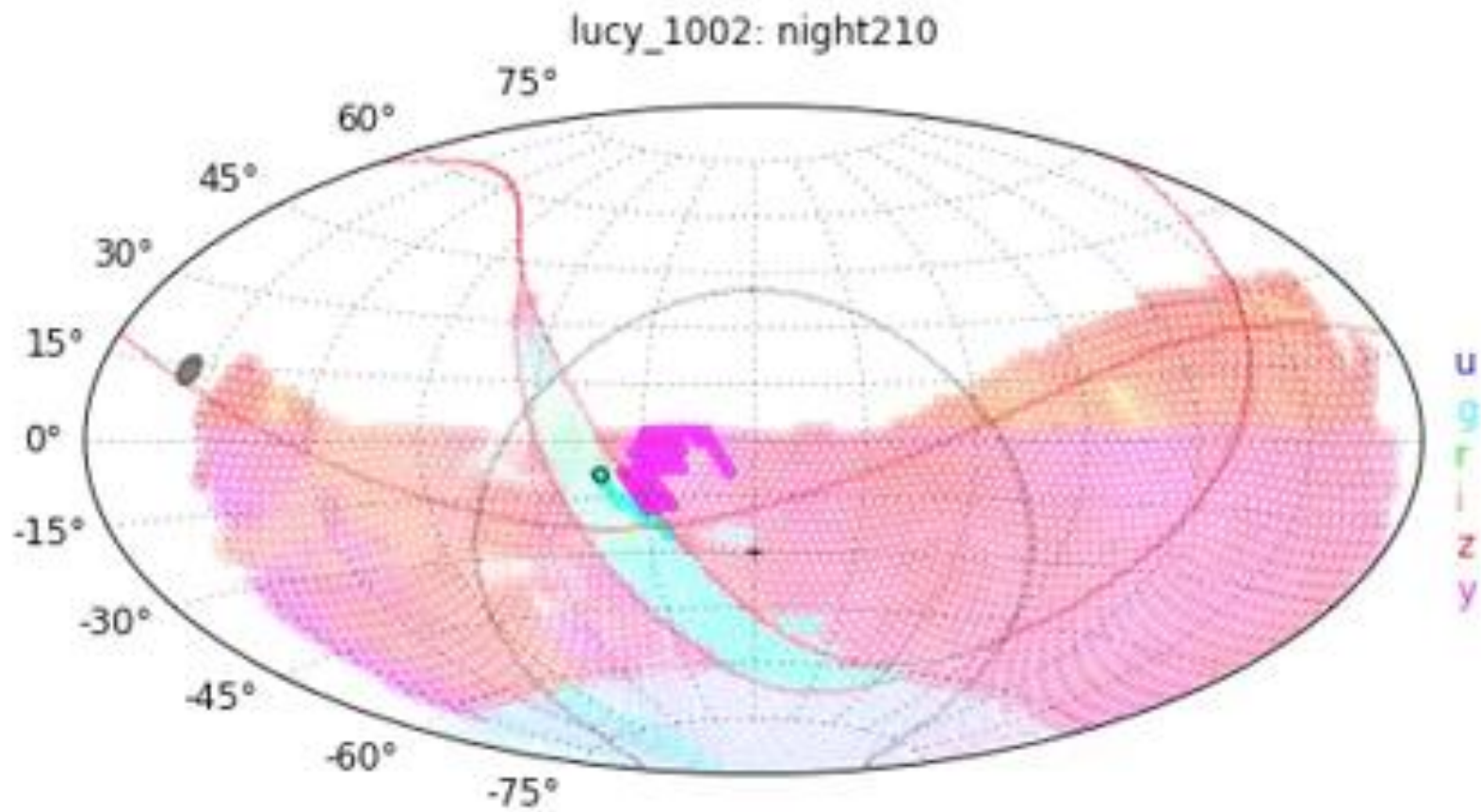
~10 billion alerts



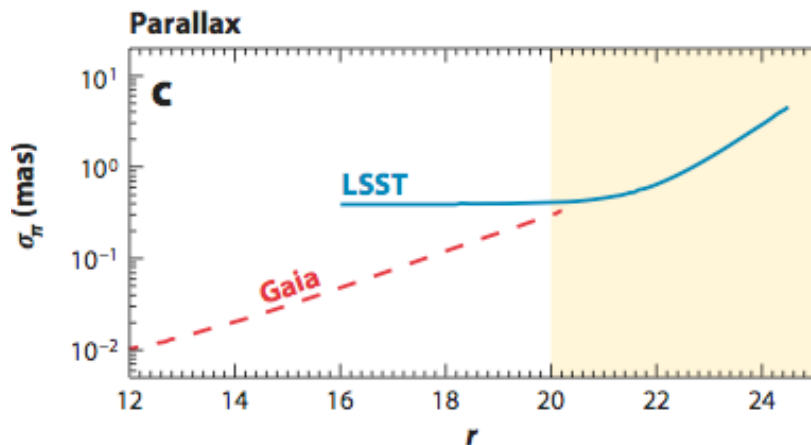
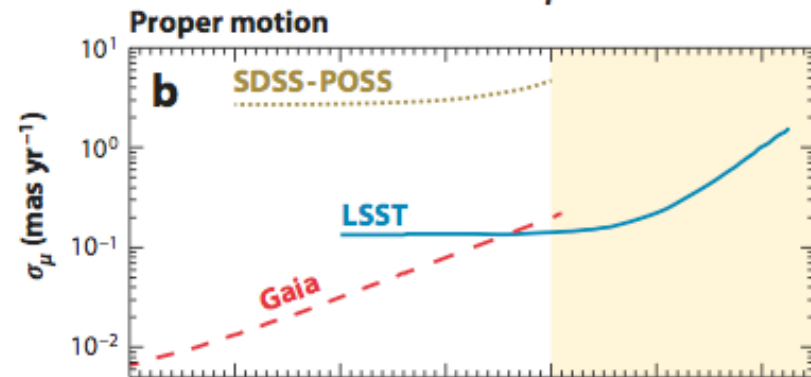
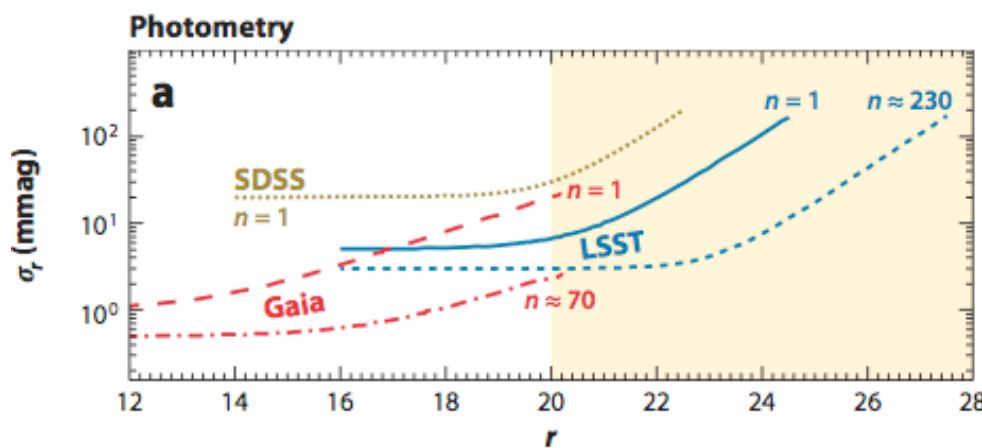
LSST Science Requirements Document ([ls.st/srd](https://www.lsst.org/lsst/srd)):
Report alerts within 60 sec after closing the shutter.

Automated scheduling of LSST observations (speed 1000x)

Time: 49562.988731



Astrometric and photometric performance comparison for Gaia and LSST

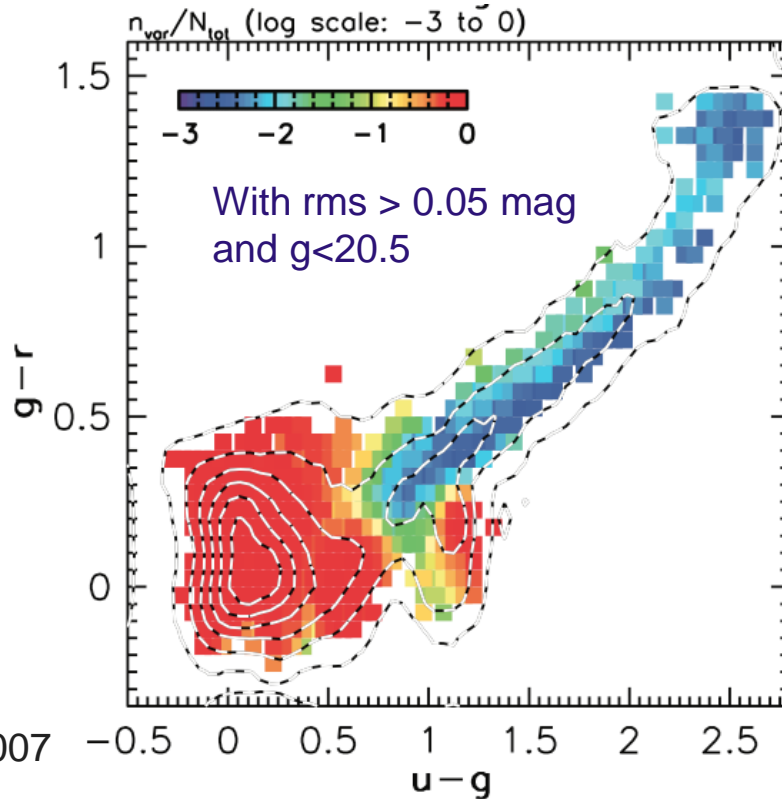


Photometric accuracy: random errors 0.005 mag, calibration to 0.01 mag; for light curves, LSST “takes over from Gaia” around $r \sim 17$

Time-resolved measurements: photometric variability, and parallax and proper motions from astrometric measurements

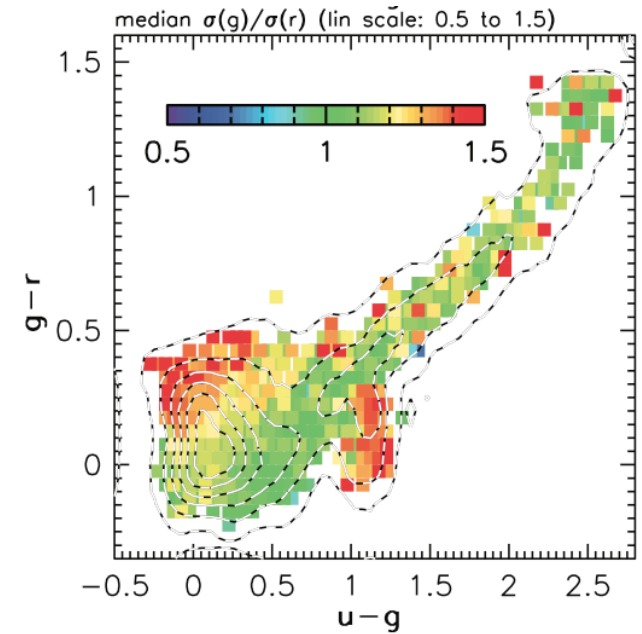
Gaia vs. LSST: complementarity of the two surveys: photometric, proper motion and trigonometric parallax errors are similar around $r=20$

The fraction of variable point sources in SDSS



Sesar+2007

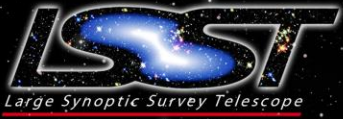
The sample is dominated by quasars and variable stars called RR Lyrae.



Rubin Observatory Construction Status



Rubin Observatory Team, August 2022, Tucson, AZ



LSST Operations: Sites & Data Flows



and a British site, too!

French Site

Satellite Processing Center
Data Release Production
Long-term Storage (copy 3)

Archive Site

Archive Center
Alert Production
Data Release Production
Calibration Products Production
EPO Infrastructure
Long-term Storage (copy 2)
Data Access Center
Data Access and User Services

HQ Site

Science Operations
Observatory Management
Education & Public Outreach

Base Site

Base Center
Long-term storage (copy 1)
Data Access Center
Data Access & User Services

Summit Site

Telescope & Camera
Data Acquisition
Crosstalk Correction

Google

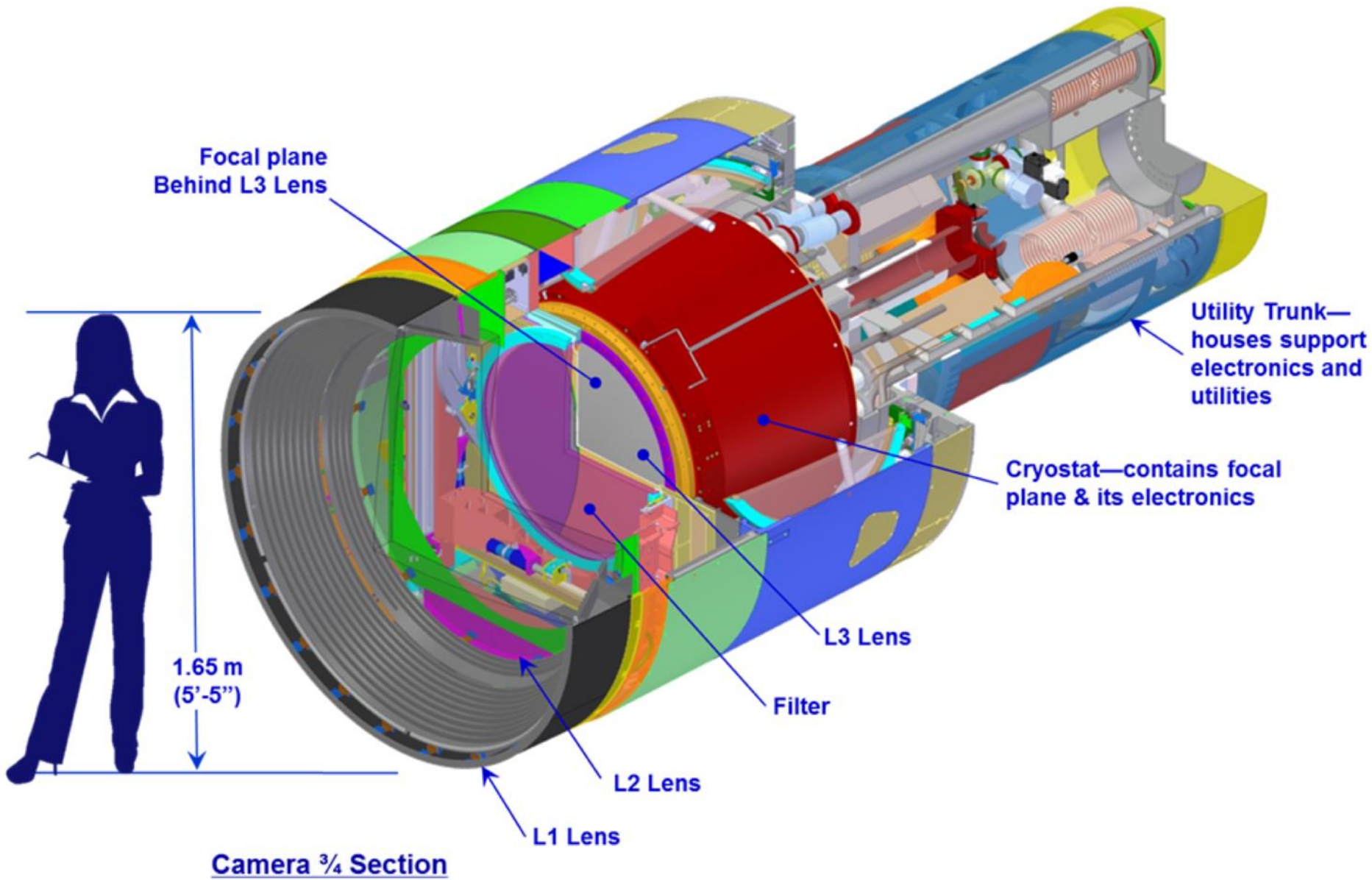
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TMA Moves

December 2022



Focal plane Behind L3 Lens

Utility Trunk—houses support electronics and utilities

Cryostat—contains focal plane & its electronics

L3 Lens

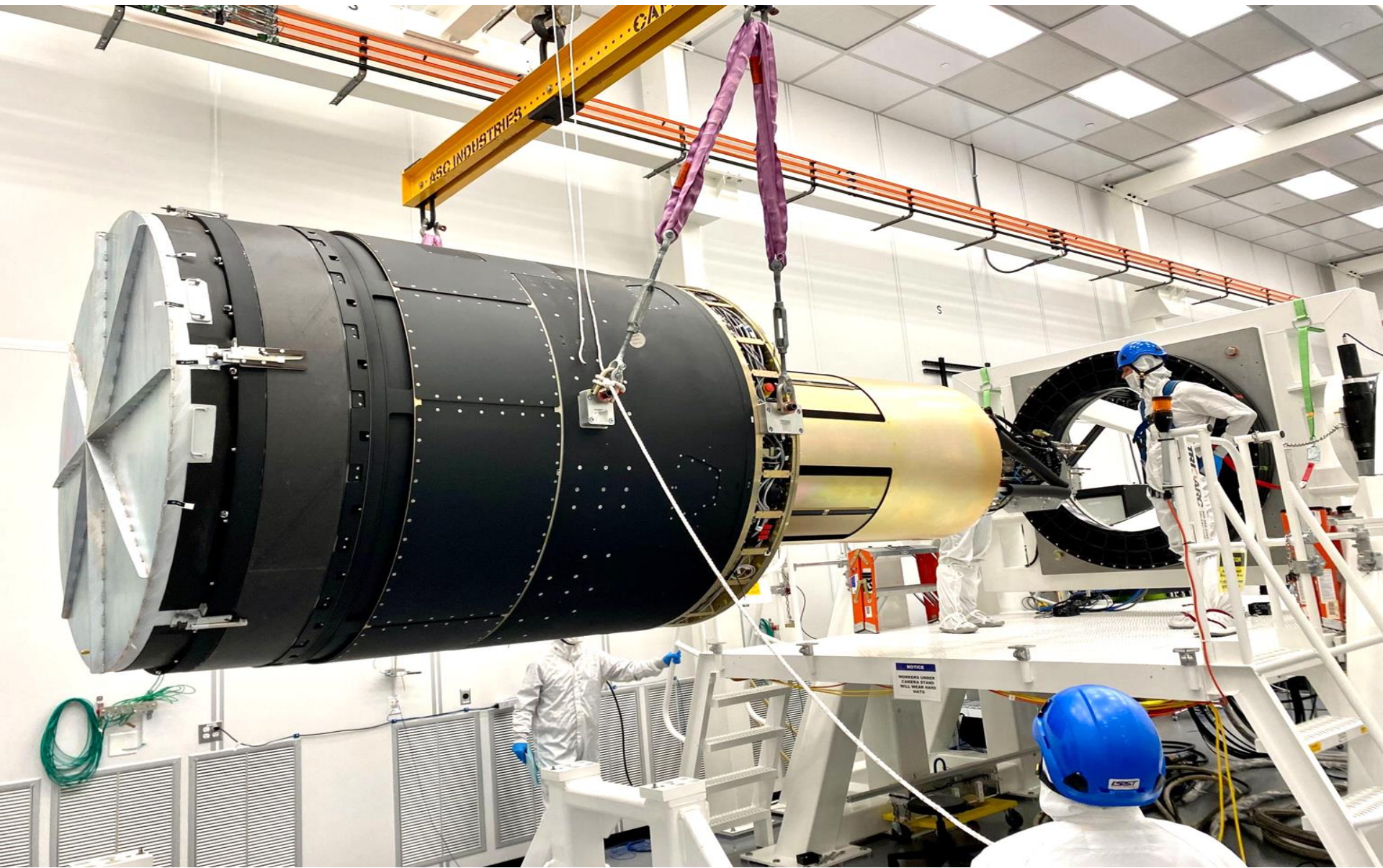
Filter

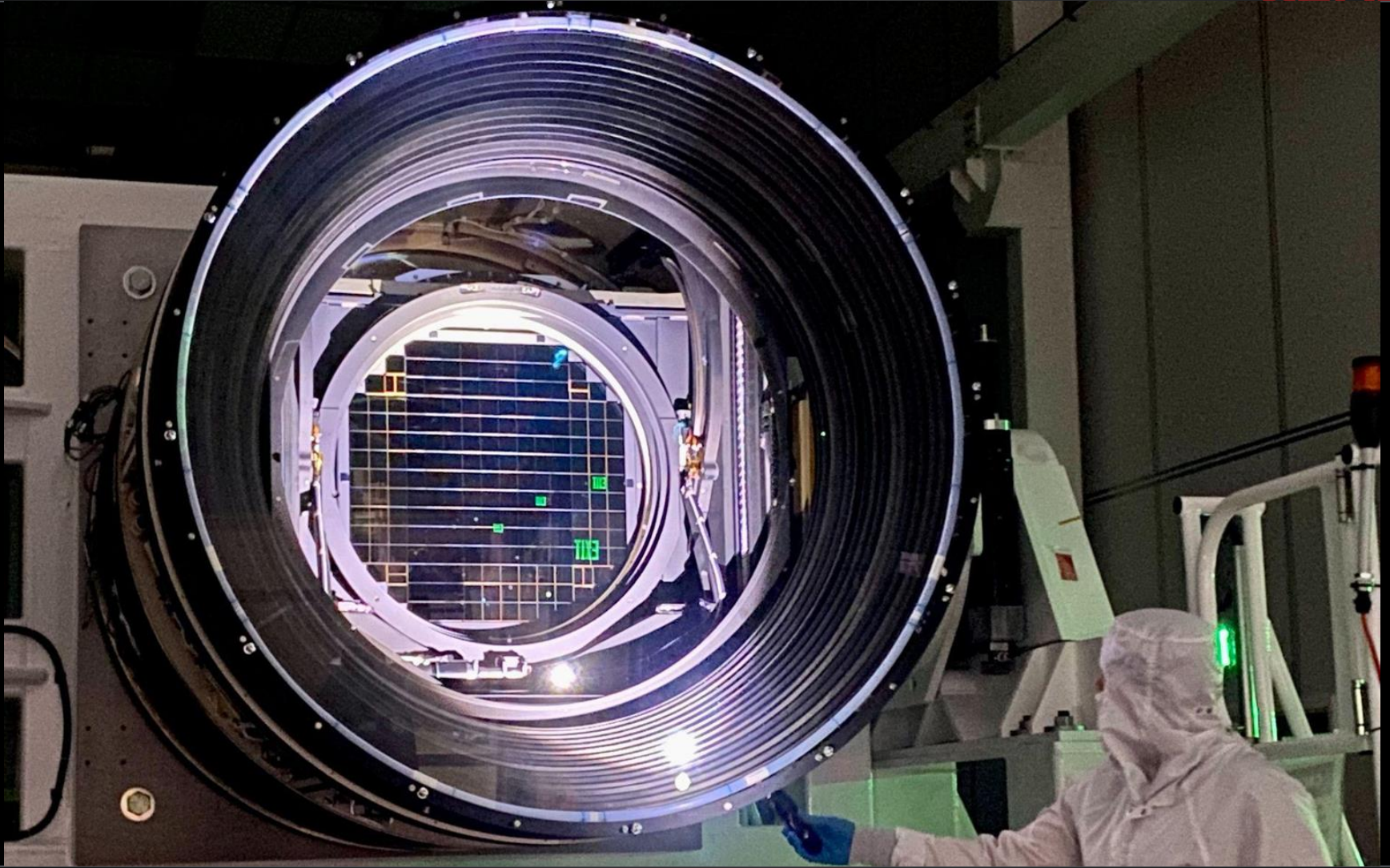
L2 Lens

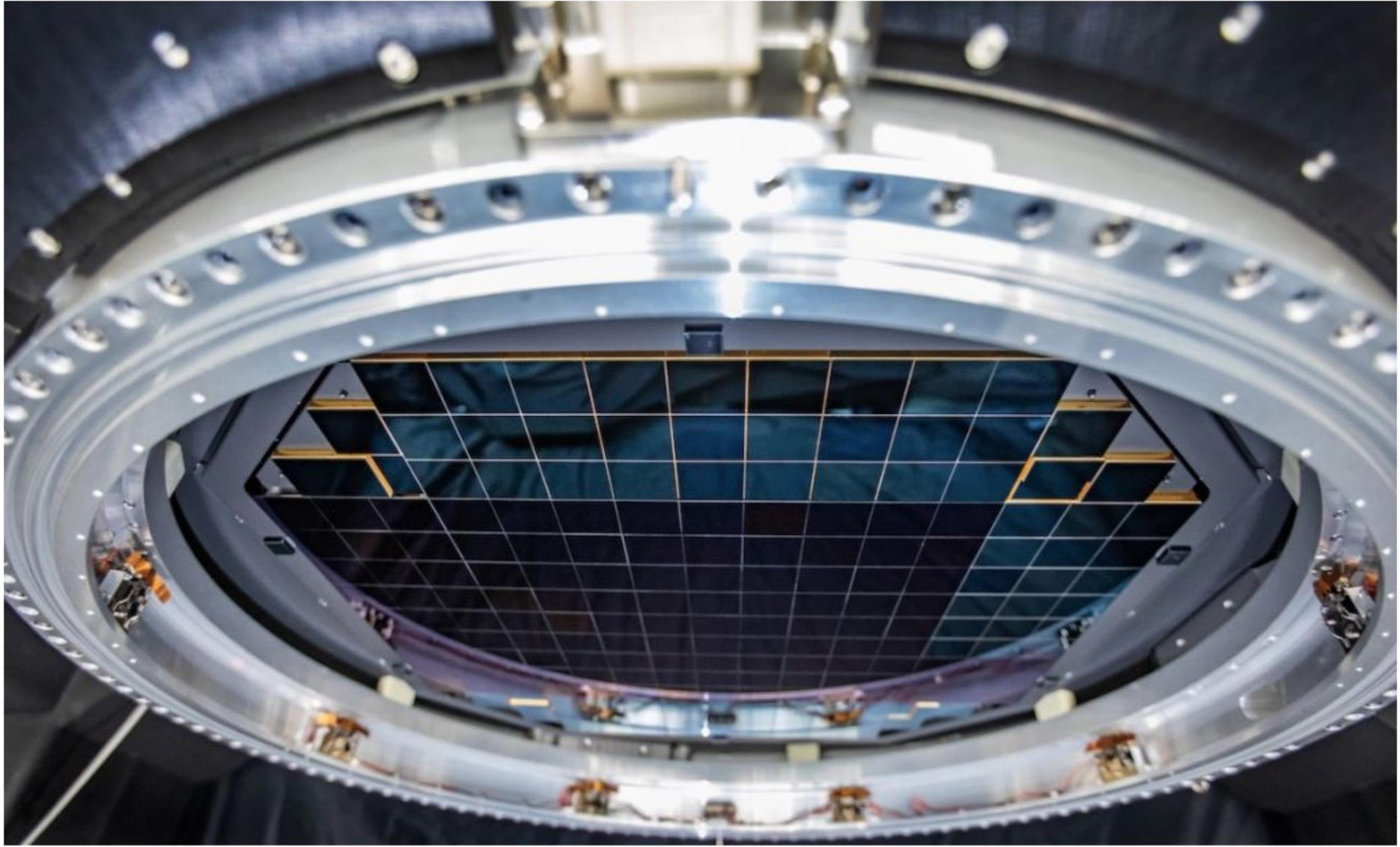
L1 Lens

1.65 m (5'-5")

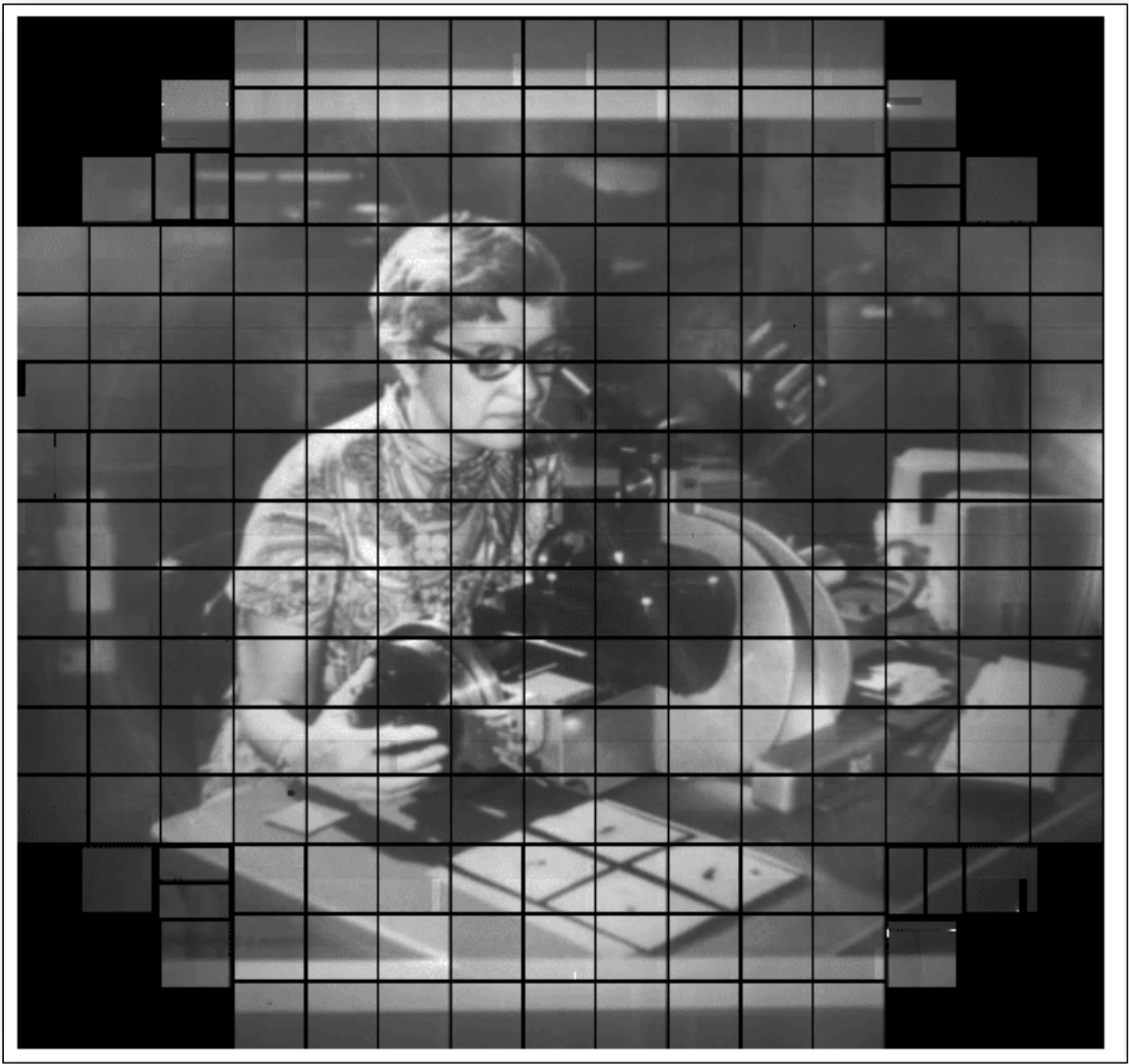
Camera 3/4 Section





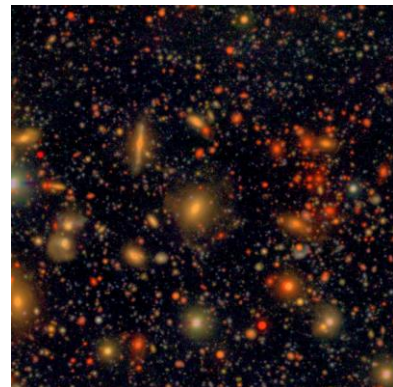
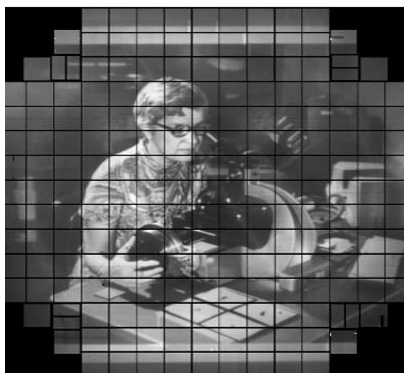
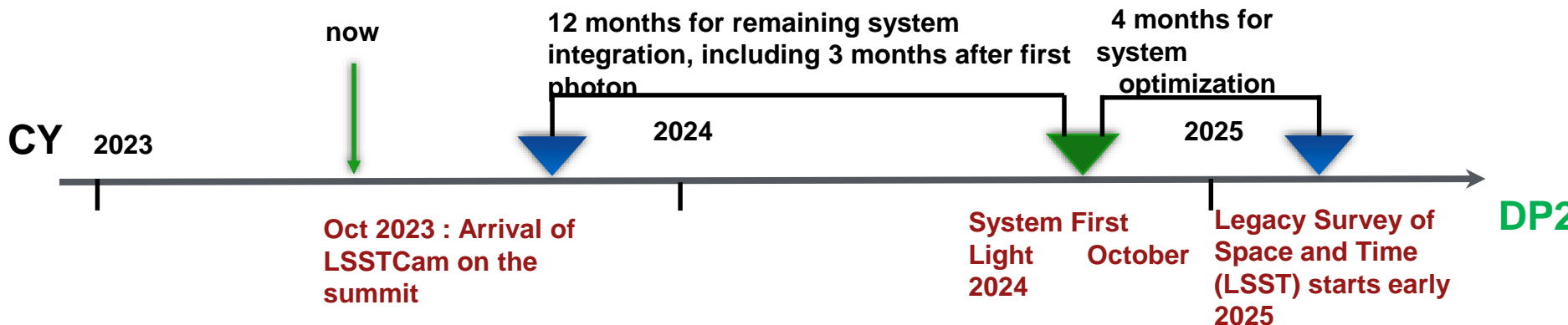


The complete focal plane of the future LSST Camera is more than 2 feet wide and contains 189 individual sensors that will produce 3,200-megapixel images.



Rubin Construction Timeline

start of LSST: early '25



Significant reduction of schedule uncertainty with the arrival of LSST Camera to the Observatory site in Chile (October 2023).

Data Preview 2 (DP2):
2 months of commissioning survey data (Aug '25) via Rubin Science Platform

Legacy Survey of Space and Time: a 10-year survey starting in 2024

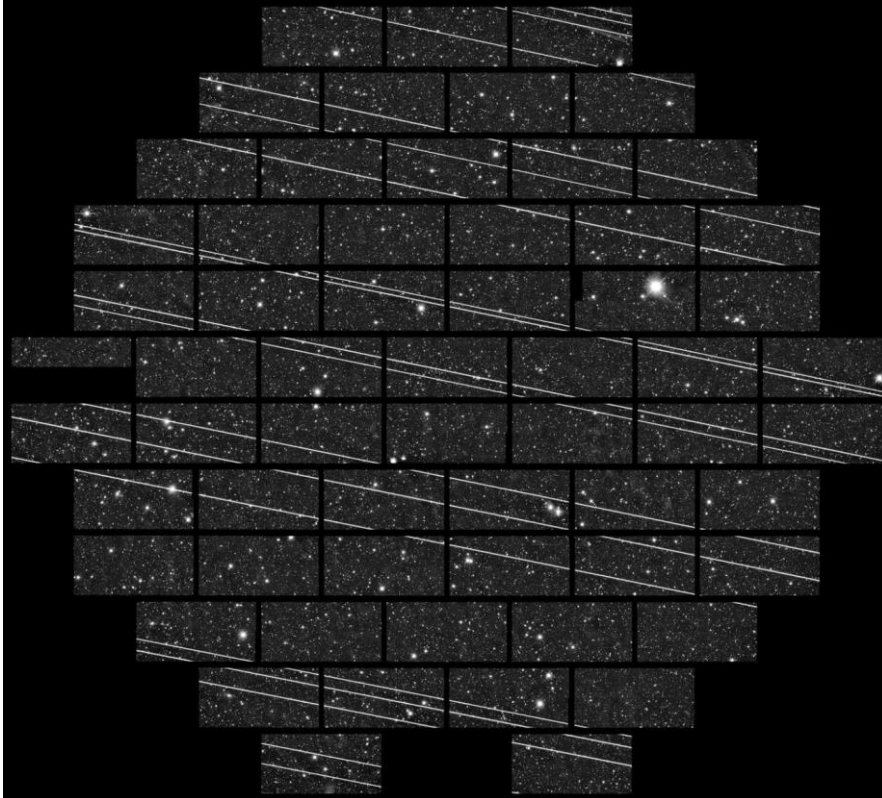
More details:
ls.st/lop

multi-color time-resolved faint sky map

- 20 billion galaxies
- 20 billion stars
- 10 billion alerts
- “millions and millions” of SNe, quasars, asteroids...



Low-Earth orbit satellite constellations



DECam image: 333 sec exposure, 19 Starlink streaks (Clara Martínez-Vázquez and Cliff Johnson)

How bad is that?

A few points to make:

- at that time, satellites were still much closer to the observer than when in their final orbits (the so-called “at station”)
- LSST visit is ~10 times shorter and the FOV is somewhat larger: ~2-3 streaks
- there are many other quantitative details that need to be taken into account...

So, really, **how bad are these satellite constellations for LSST?**

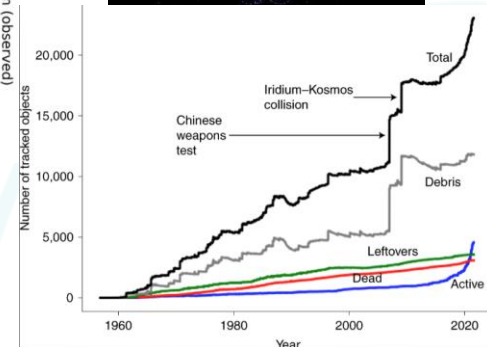
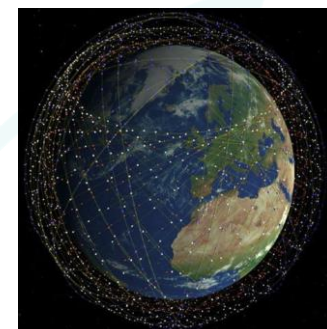
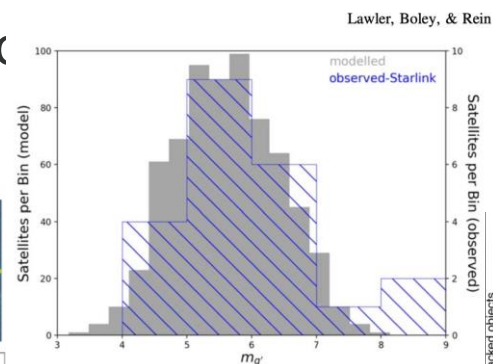
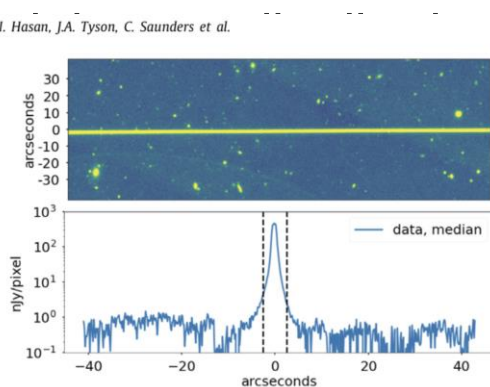
Impact of LEO Satellite Constellations on Rubin Observatory and LSST science

Quantitative assessment depends on several imperfectly known quantities:

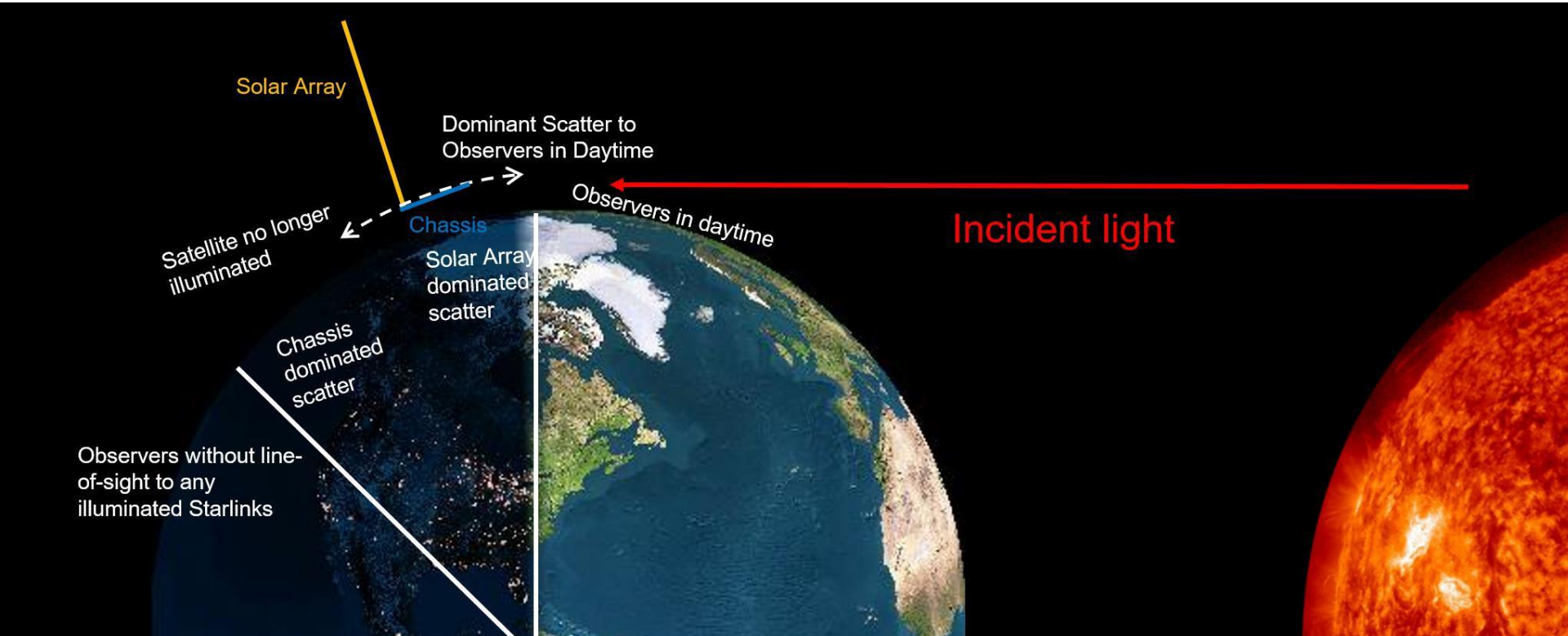
1) The number of satellites and their characteristics

2) The satellite light properties

3) Impact on LSST image quality and mitigation



2) The satellite brightness distribution



Summary (from Rubin/LSST point of view)

With tens of thousands of LEOsats, generally ***no combination of mitigations can completely avoid the impacts of the satellite trails*** on LSST science programs.

However, current predictions of the impact correspond to a “**nuisance**” that we have to plan for (~1% of pixels lost), rather than a “catastrophic” impact (>10% of pixels lost).

We need to continue to **constructively** communicate with satellite providers.



Starlink Satellites Overhead



Starlink Satellites pass overhead near Carson National Forest, New Mexico, photographed soon after launch.

Credit: M. Lewinsky/Creative Commons Attribution 2.0

A potentially **misleading** figure: most streaks are actually stars!