SOLAR OBSERVATIONS IN BULGARIA AND COLLABORATIONS IN SPACE WEATHER RESEARCH

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Outline







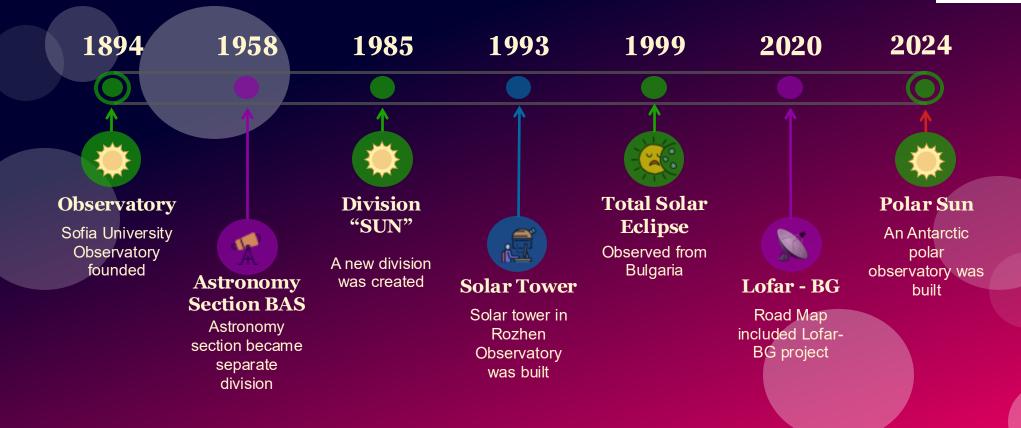






Very Brief History





Domains of Research

filaments

- Flares

Solar Activity Solar Eclipses Space Weather - New indirect index of the Faint H-alpha emission - Geoeffective active - Radio bursts in the solar corona solar activity regions - Activity of the solar - Corona polarization - SEP events - Machine learning, prominences automatic feature recognition - Distribution and large-- Rapid Changes in the - Geomagnetic storms scale motions of the solar Visible Wavelength of the

Corona

- Impact on satellites

Radio Sun

Bulgarian-Egyptian project

On space weather effects At near Earth environment - from remote observations And in situ particle forecasting to impacts on satellites (finished, 2022-2024)

Work packages

- **1. SEP events -** Estimating the fluences of the different energetic particles; SEP forecasting of selected events; Development, validation and operational application of new SEP forecasting methods
- 2. **Solar events** Remote observations and in depth data analysis by performing monitoring the origins of SEP events in optical and radio wavelengths; Data analysis and deduction of physical parameters of the solar eruption
- **3. Impact on satellites** Qualitative and quantitative evaluation of satellite risks caused by space weather events; Investigating the different effects of hazardous energetic particles on a GEO spacecraft

Publications:

Universe 2023, 9(4), 179 (15pp.) https://doi.org/10.3390/universe9040179

Astron. Soc. "Rudjer Bošković" No 25, 2023, 125-135 https://astro.bas.bg/conf_proc/book_XIIIBSAC.pdf

Atmosphere, 2023, 14(12) https://doi.org/10.3390/atmos14121744

Astronomy, 2023, 2(3), 165-179 https://doi.org/10.3390/astronomy2030012

Proceedings SES 2023, 67-70

http://space.bas.bg/SES/archive/SES%202023_DOKLADI/ ROCEEDINGS%20SES%202023.pdf

Bulgarian-Serbian project

Active Events on the Sun.Catalogs of Proton Events and Electron Signatures In X-Ray, UV and

Radio Diapason.

(2023-2025)

Bg-PI: M. Dechev

Rs-PI: Z. Simic



AIM PROTON EVENTS

SXR FLARES

RADIO BURSTS

GEOMAGNETIC STORMS

TYPE II BURSTS

The Catalog list the proton events from SOHO/ERNE instrument presented separatly in solar cycles SC23 (1996-2008) and SC 24 (2009-2017).

The Catalog utilizes the highest tempporal resolutionas provided, namely 1 min.

Results from the final SEP catalog are available here:

Miteva et al. Atmosphere (2024)

Preliminary results from the catalog are published here:

Miteva et al. SES-Proceedings (2023)
Miteva et al., Bulgarian Astronomical Journal, Vol. 33, pp. 99-108 (2020)

Joint Observations and Investigations of Solar Chromospheric and Coronal Activity (2023-2025)

Bg-PI: R. Miteva At-PI: W. Poetzi

1. Parameter study of geoeffective active regions (completed, 2024)

https://doi.org/10.3390/atmos15080930

2. Confined M-class flares in solar cycles 23&24 (completed, 2025)

https://astro.bas.bg/AIJ/issues/n43/RMiteva.pdf



AIM

В

To set up the Rozhen Chromospheric Telescope (RCT), and develop standardized solar observing methodology and products, complementary to the Kanzelhohe Patrol Instrument (KPI) by means of strong technical cooperation between the team members.

To carry out combined solar observations with the two instrument suites and external (freely available space-based) resources, in order to study chromospheric signatures of quiet sun and pre-eruptive active regions and multi-wavelength manifestation of solar eruptive phenomena, their morphology and kinematics.

Work Package #1

Technical support of NAO-Rozhen Chromosphere Telescope and observation campaigns with KSO facilities

- · Task 1.1: Telescope installation
- Task 1.2: Data processing
- · Task 1.3: Observation Campaign
- Task 1.4: Image enhancement

Work Package #2

Joint investigations of solar chromospheric and coronal activity

- Task 2.1: Chromospheric Signatures of Quiet Sun and Pre-Eruptive Configurations
- Task 2.2: Multi-wavelength study of solar activity phenomena, their morphology and kinematics

Work Package #3

Dissemination of the project results

- . Task 3.1: Project web-site
- Task 3.2: Scientific dissemination



Observatory Kanzelhöhe in Austria

Long tradition in Solar observations (since 1944)

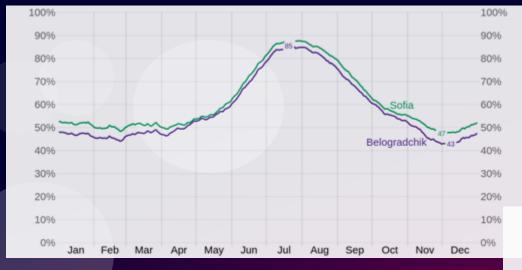
Very experienced staff (scientific and technical)

Long data archive

365(6) days per year



Some Weather Considerations



Up to 20% clouds - clear

The clearer part of the year in Belogradchik begins around June 7 and lasts for 3.5 months, ending around September 23.





Instrumentation: Telescope - Belogradchik Observatory

Celestron Schmidt-Cassegrain telescope SC 279/2800 C11 OTA

For filter f/d = 1/30 is needed \Rightarrow

- 1) enlarge focal length;
- 2) reduce aperture;

Reduce aperture advantages:

- less heat;
- less light camera exposure still very short;
- cheaper optical components;
- f/l for filter is better;
- resolution still largeenough for seeing conditions;



Instrumentation: Filter, Camera - Belogradchik Observatory



Hydrogen Alpha Quantum 1.25 DayStar

(FWHM) < 0.5Å filter's wavelength output (accurate to 0.1Å) and red /blue wing shift buttons Temperature stabilized



Common Vision Spark Series, SP-12401M-USB

Light Spectrum Visible + NIR Resolution Resolution WxH Frame rate

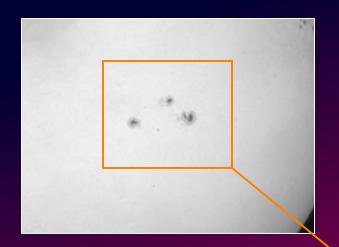
12.4 MP 4112 x 3008 px

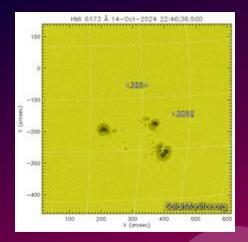
23 fps

ACAG9 October 13–16, 2025 · Helwan, Cairo, Egypt (NRIAG)

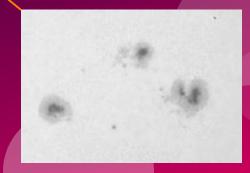
Overexposed image shows prominence at limb. Structures are visible. Relatively good contrast.

First tests – Belogradchik Observatory

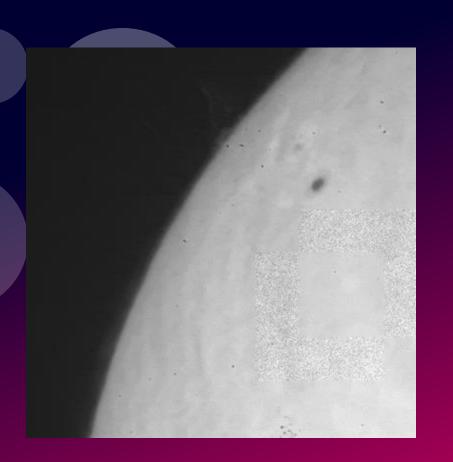




Solar spot group on 2024-10-11 and the same group in Solar Monitor. Resolution is sufficient to resolve details. Seeing conditions were bad (wiggly limb), but umbra and penumbra well separated



2024-10-14



2025-05-14

X2.7 flare

First tests – Shumen University Observatory



Instrumentation and test image on 2025-09-23

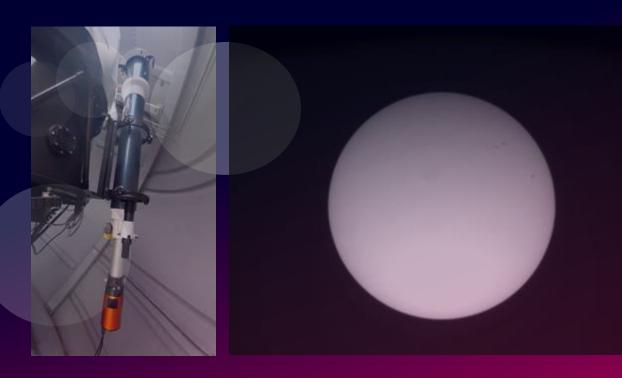
The second station, AO-Shumen University, is located on the Shumen plateau (at about 500 m elevation). The instrumentation used (Fig. 2, left) is as follows:

- Telescope: Coronado ST 40/400 PST Personal Solar Telescope OTA, D = 40 mm, focal ratio: f/10
- Mount/guiding: azimuthal, 'HelioFind' automatic Sun-guiding system
 - Camera: NexImage 5 CCD camera
 - Filter: integrated H-alpha

The image was taken after a three-fold extension of the focal length to 1200 mm using a projection eyepiece.

The image also shows a weak, but recognizable interference pattern, with numerous sunspot groups.

First tests – Shumen St. George school Observatory



Instrumentation and test image from St. George school and preschool observatory on 2025-06-04.

The Instrumentation to be used at St. George School and Preschool-Observatory:

- Guiding telescope: Refractor
 Celestron Omni XLT 120, D = 120 mm,
 f = 1000 mm, focal ratio: f/8.3
 - Mount/guiding: 10Micron
- Camera: Atik Horizon Color CMOS Panasonic MN34230, chip size: 17.6 x 13.3 mm (22 mm diagonal), pixel size: 3.8 µm, pixel area: 4644 x 3506
- Filter: various metallized filters, glass and special film, for example AstroSolar Photo Film OD 5 from Baader

Several sunspot groups are visible at the western limb. According to the Solar Monitor database these are ARs 14100 (upper) and 14099 (lower)

Lofar-BG



LOFAR is a multi-functional, highly innovative, pan-European distributed low-frequency radio telescope, operating between 10 and 240 MHz. It was developed by the Dutch Institute for Radio Astronomy (ASTRON) with the goal of exploring the early, distant Universe, solar activity, and the terrestrial atmosphere.



Top left: the Superterp core of LOFAR. Top right: The Chilbolton UK LOFAR station. Bottom: a core station with schematics of the HBA and LBA antennas. Credit: ASTRON





BG Polar Astronomical Observatory

The Influence of Solar Activity on Antarctic Ionospheric Dynamics and High-Energy Particle Fluxes





- 1. Study of the effects of solar flares on the ionospheric D-layer using VLF observations of distant terrestrial sources.
- 2. Study of the properties of Type II and Type III solar radio bursts at HF radio frequencies between 75-750 MHz.
- 3. Observations of secondary cosmic ray fluxes in the area of the Bulgarian Antarctic Base using particle detectors.
- 4. Observations of disturbances in the Earth's magnetic field in the area of the Bulgarian Antarctic Base, caused by solar eruptions in interplanetary space.

Acknowledgements

Bilateral project: Bulgaria-Austria https://astro.bas.bg/project-sun/

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