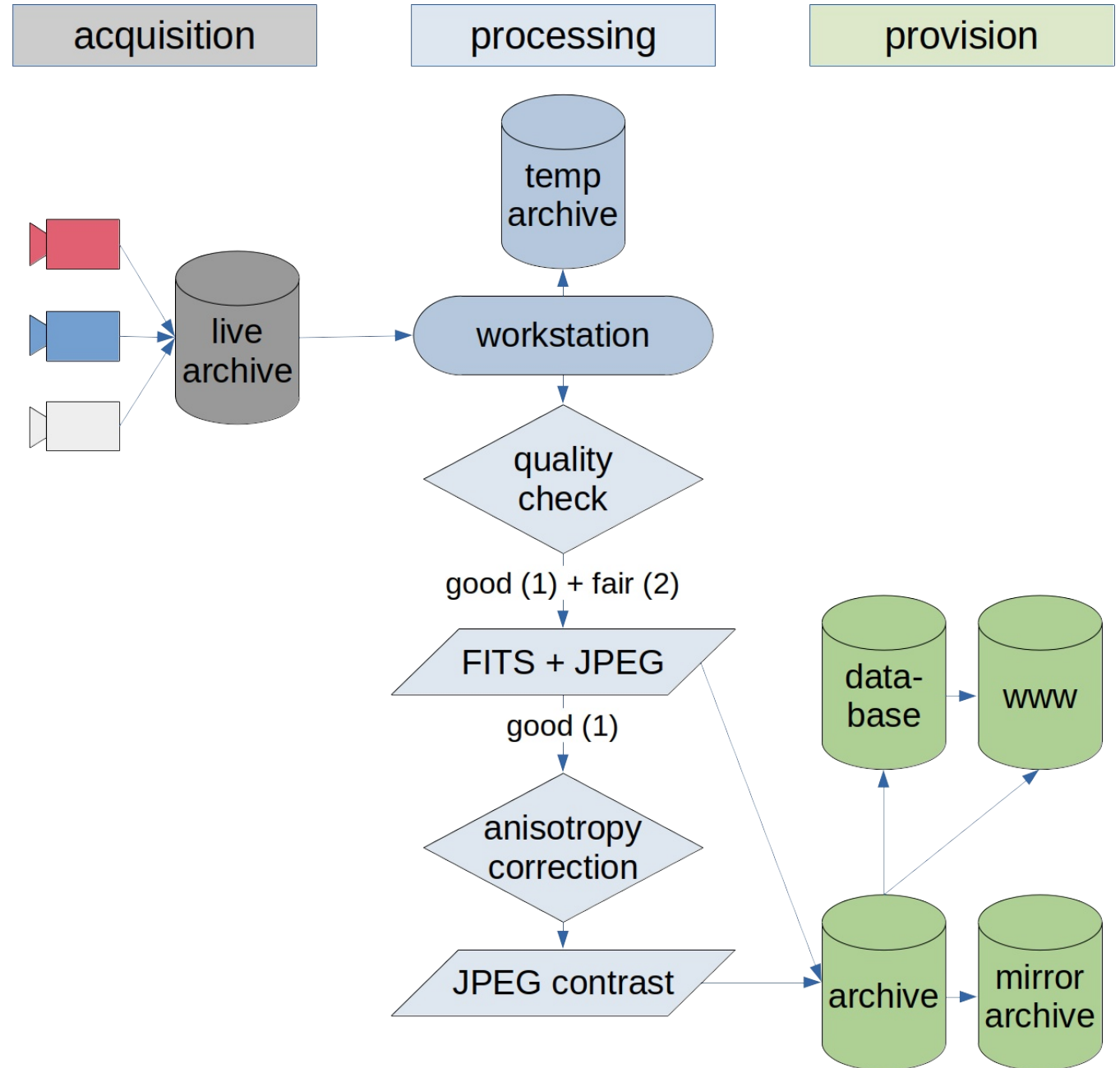


overview of  
our system

3 branches  
=  
3 separate  
systems

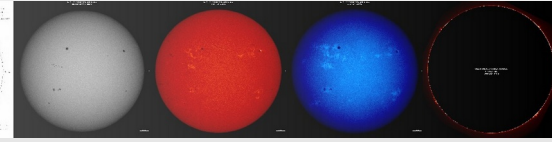
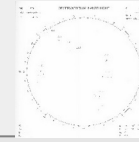
(H-alpha has a  
4<sup>th</sup> branch for  
flare detection)



## The observation tower

- outside and inside white
- ventilation removing air at lowest part of the tower



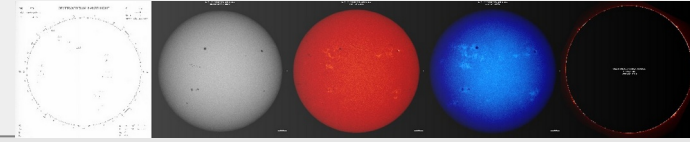


## The socket

- the telescope mounting is seated on a concrete tower
- no contact to the floor and dome

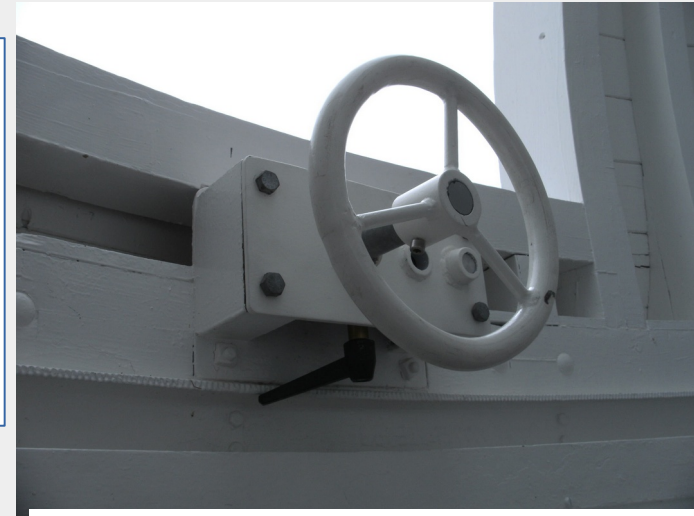




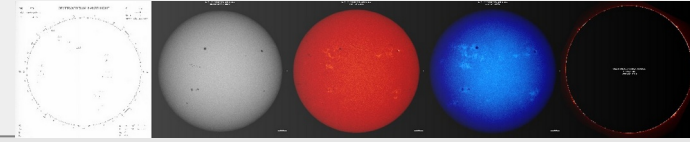


# The dome

- the slit is opened/closed manually (we are there, so no need of loose cables hanging around)
- the dome moves with the telescope – light sensors detect slit edges (only one direction!)

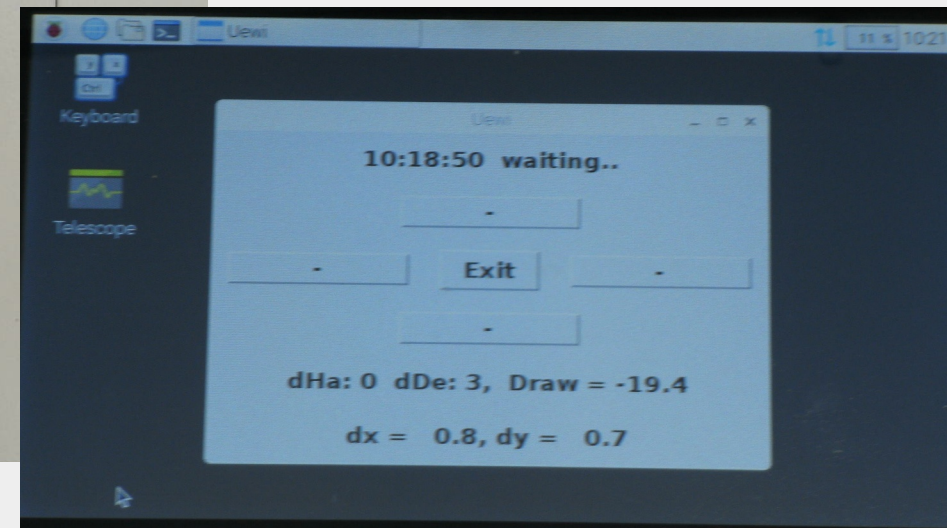
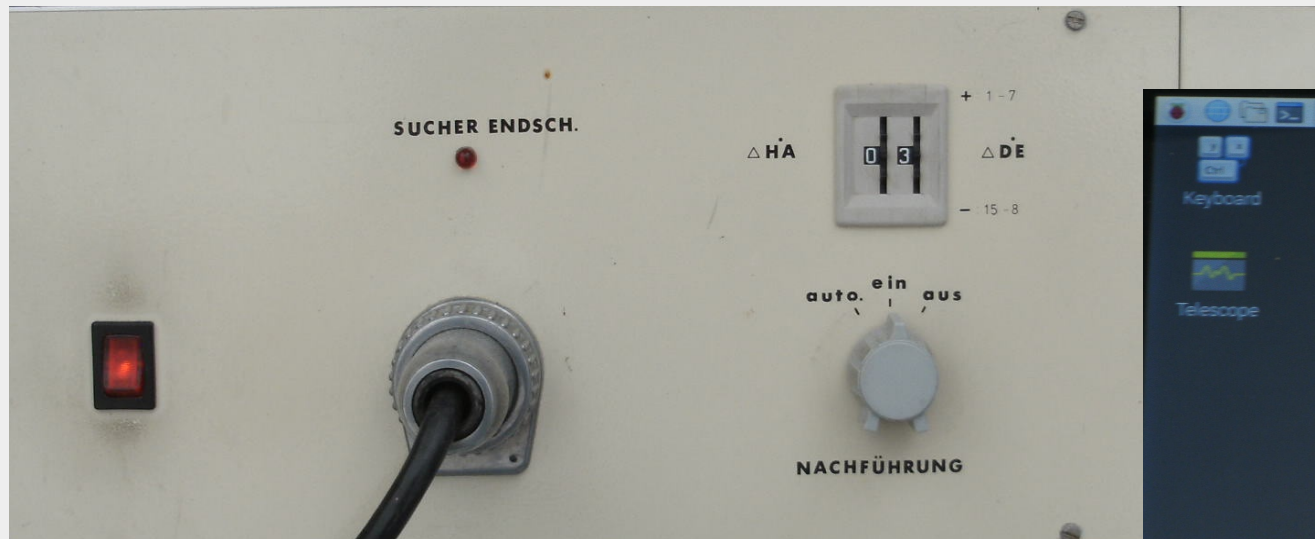


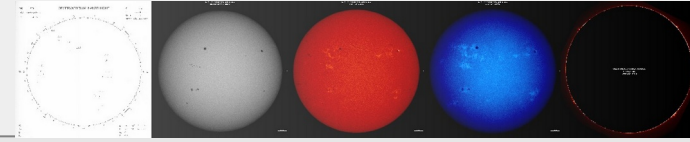




# The telescope guiding

- the principal guiding is done via an old but robust microprocessor system
- this guiding is not perfect, it has some switches to adapt it to different vertical and horizontal speeds (seasonal effects of Sun motion in sky)
- for exact guiding we send via a serial interface correction codes obtained from solar disc positions of our observing cameras (Raspberry Pi)





# The mount

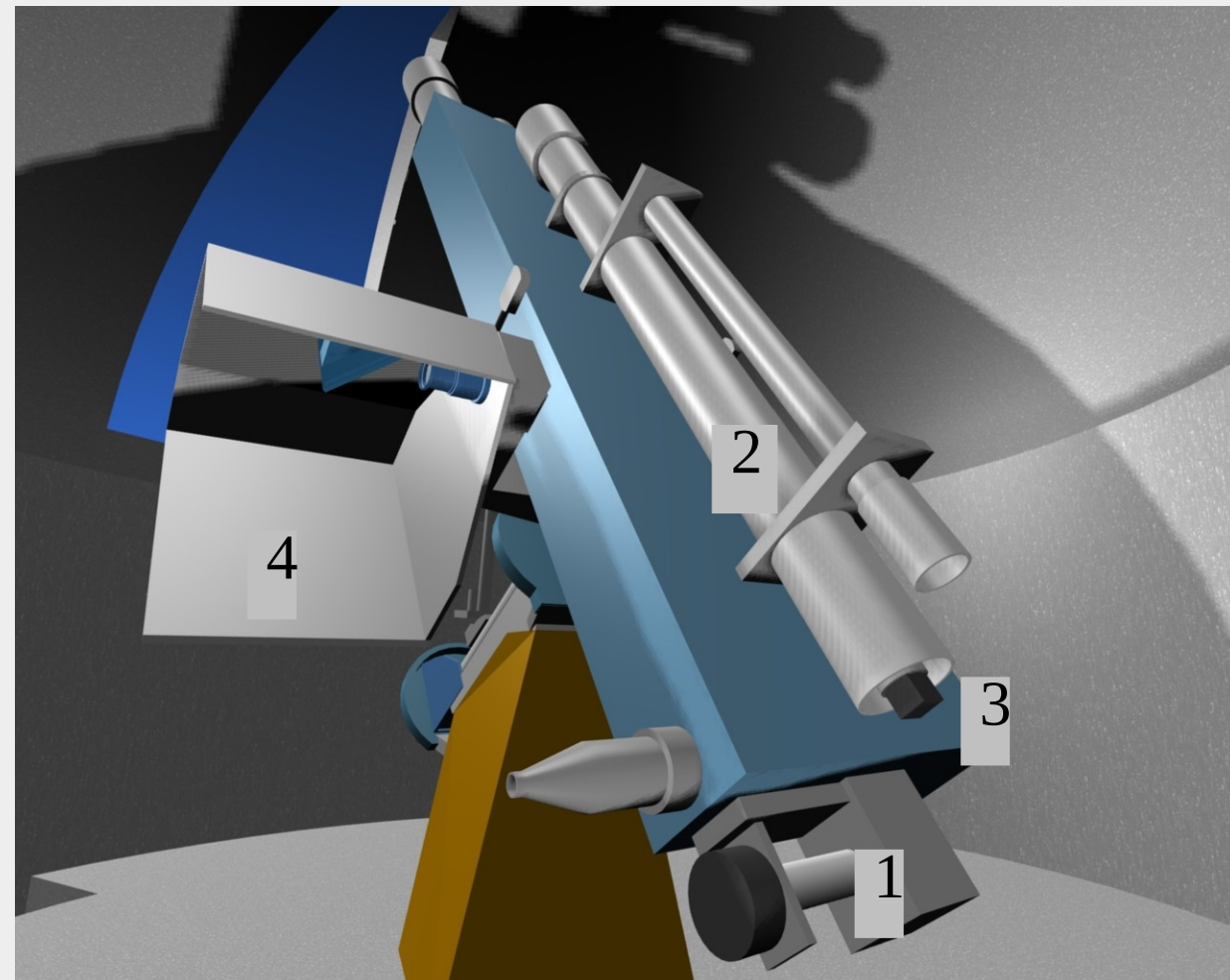


- open fork  
equatorial mount
- very stable
  - easy to control
  - balanced

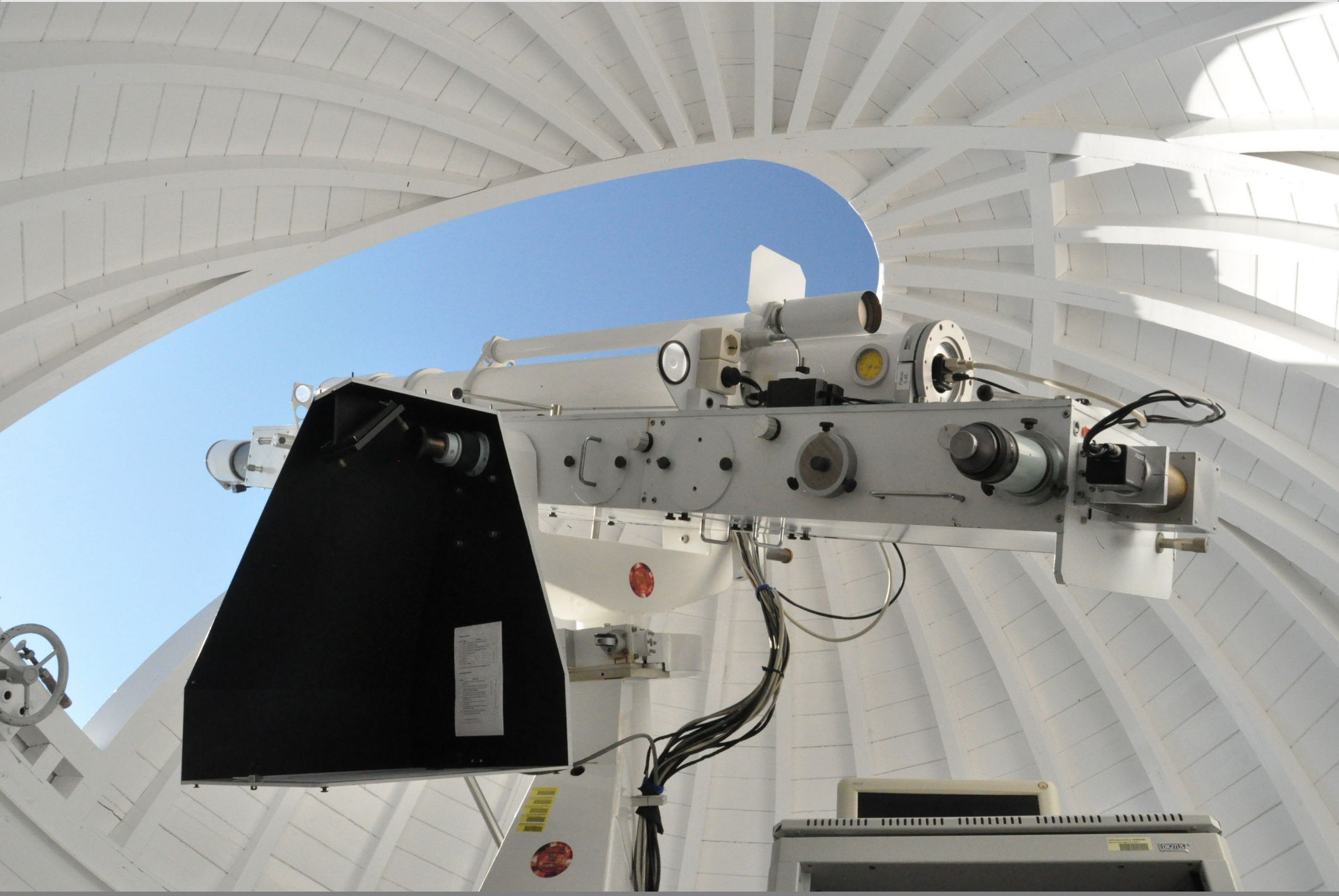
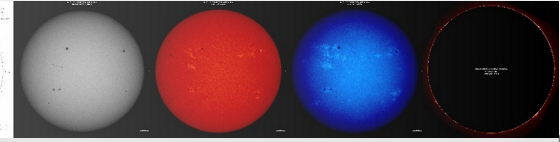
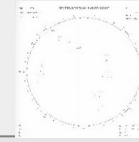


# The telescope

- 1) H-Alpha
- 2) Phoka (White-light)
- 3) CaIIK
- 4) Drawing device



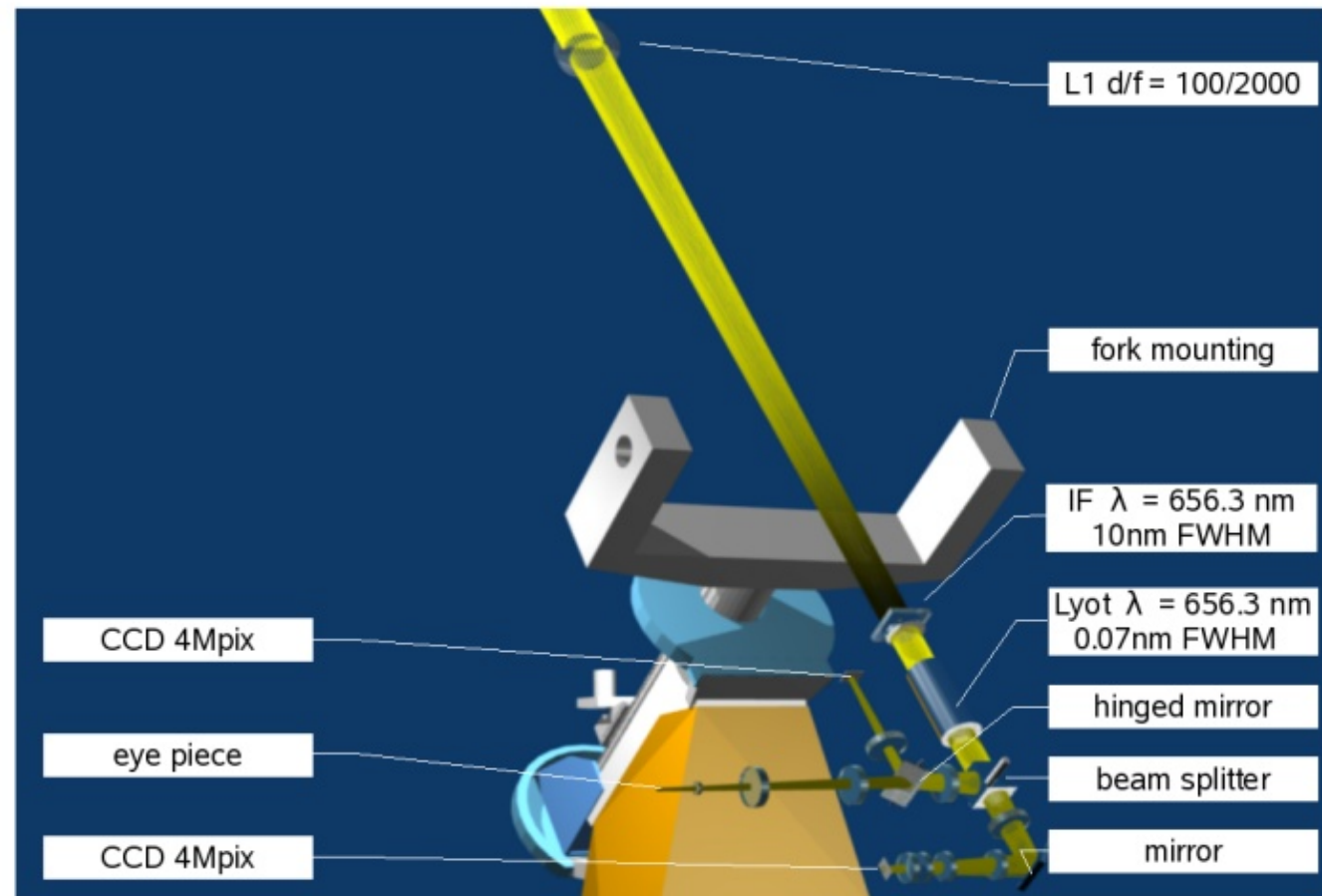
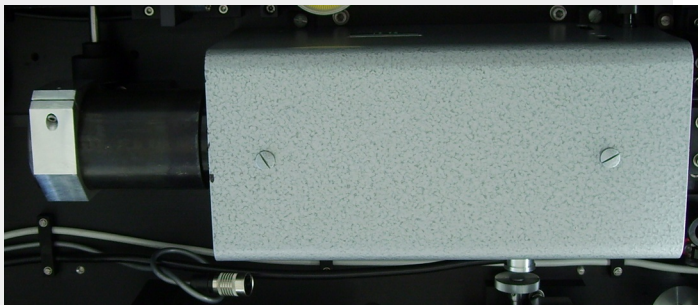






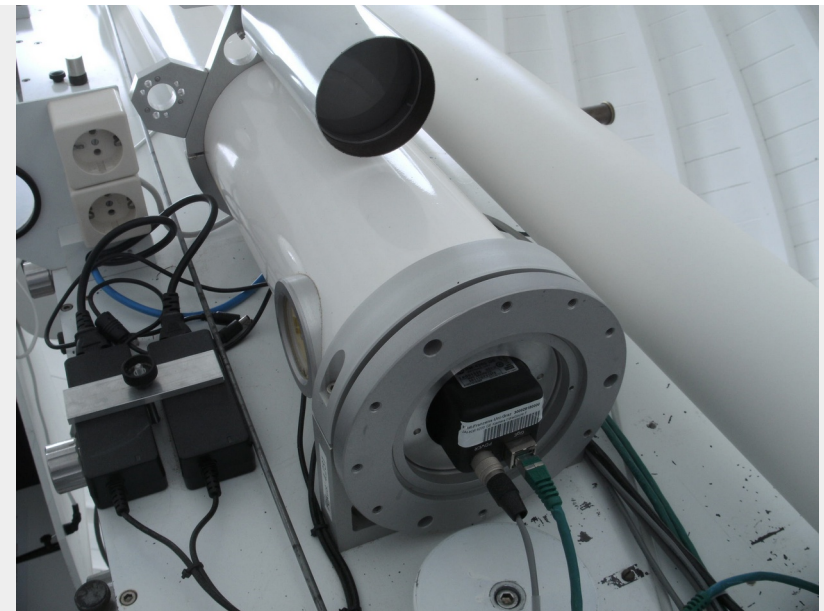
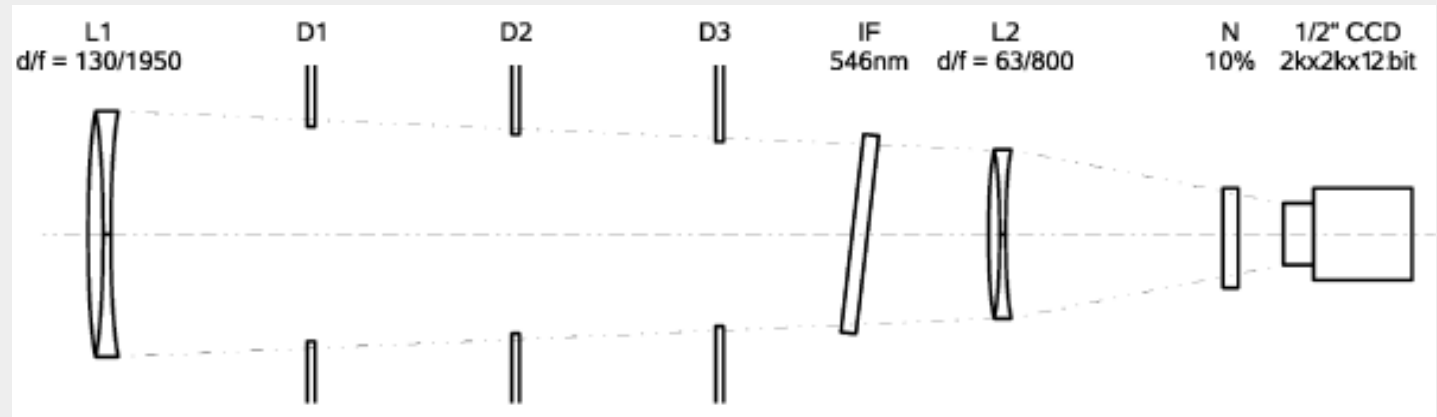
# H-alpha telescope

- prefilter (10nm)
- Lyot filter (0.7Å)
- 2 cameras (main, high speed)
- eye piece



# White-light telescope (phoka)

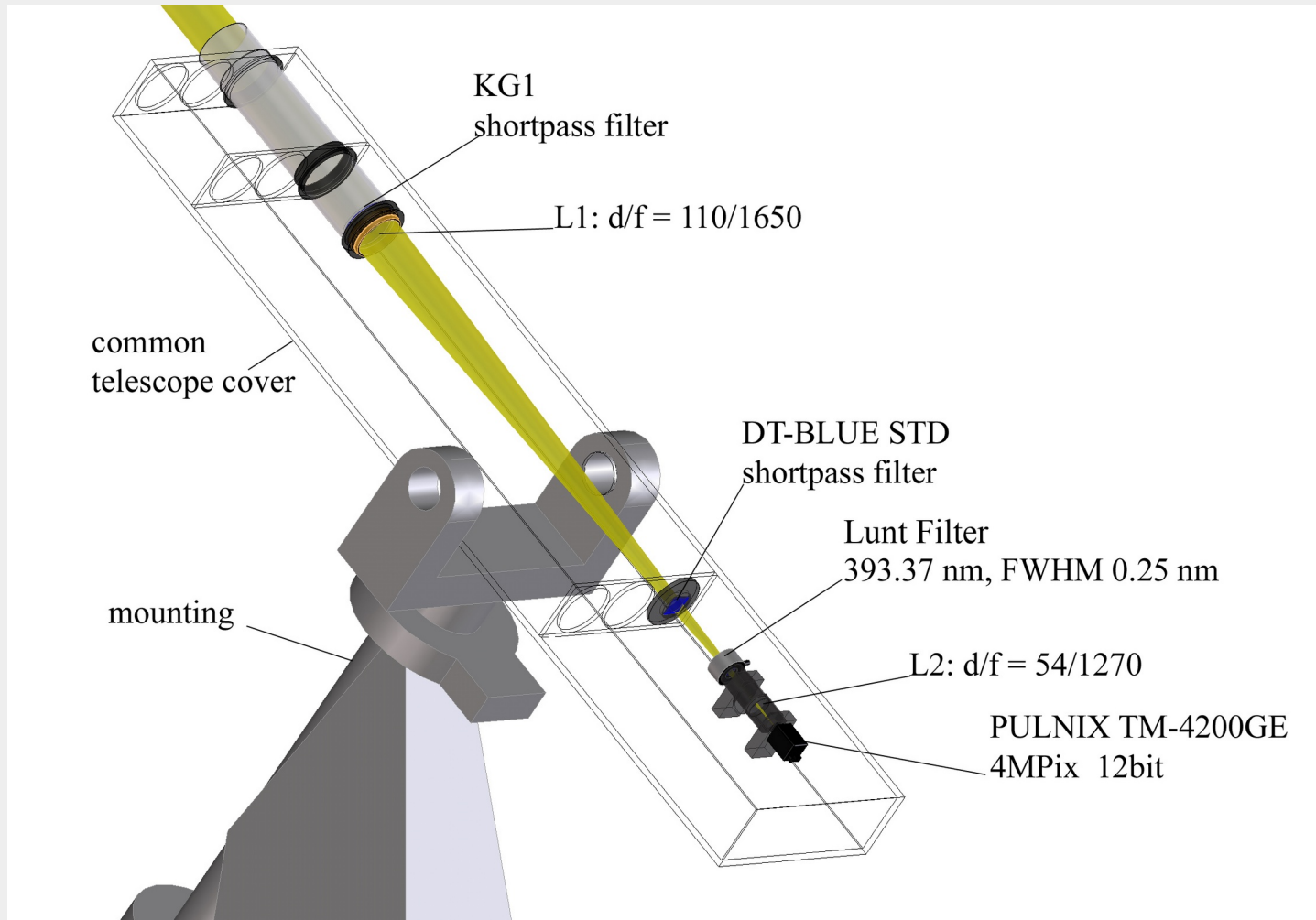
green filter  
(546nm/10nm)  
gold coating on  
objective lense



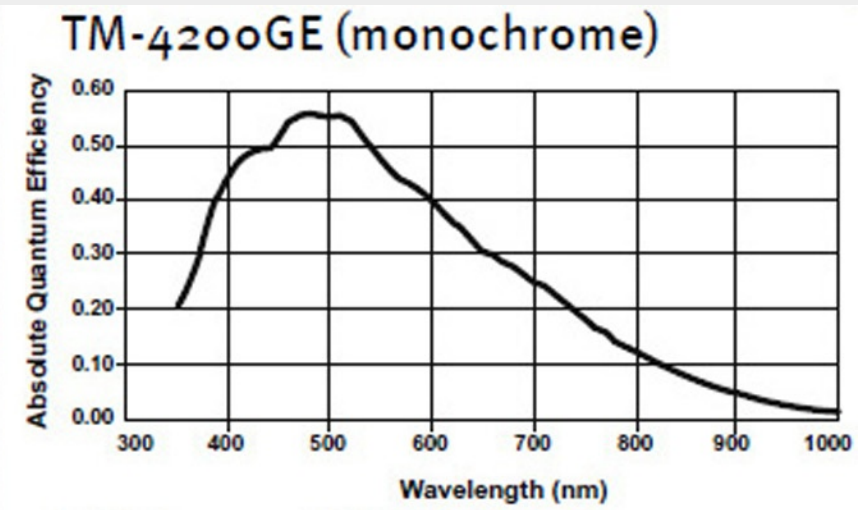


# Call K telescope

Lunt filter (cheap!)  
2.5Å

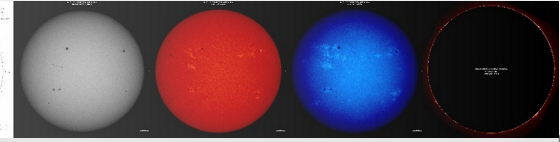
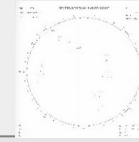


# Cameras



- same camera in all systems
- 4 MPix 12bit monochrome
- 7 images/sec
- Gigabit interface
- exposure times 1.5 to 35ms (2 ms on average)





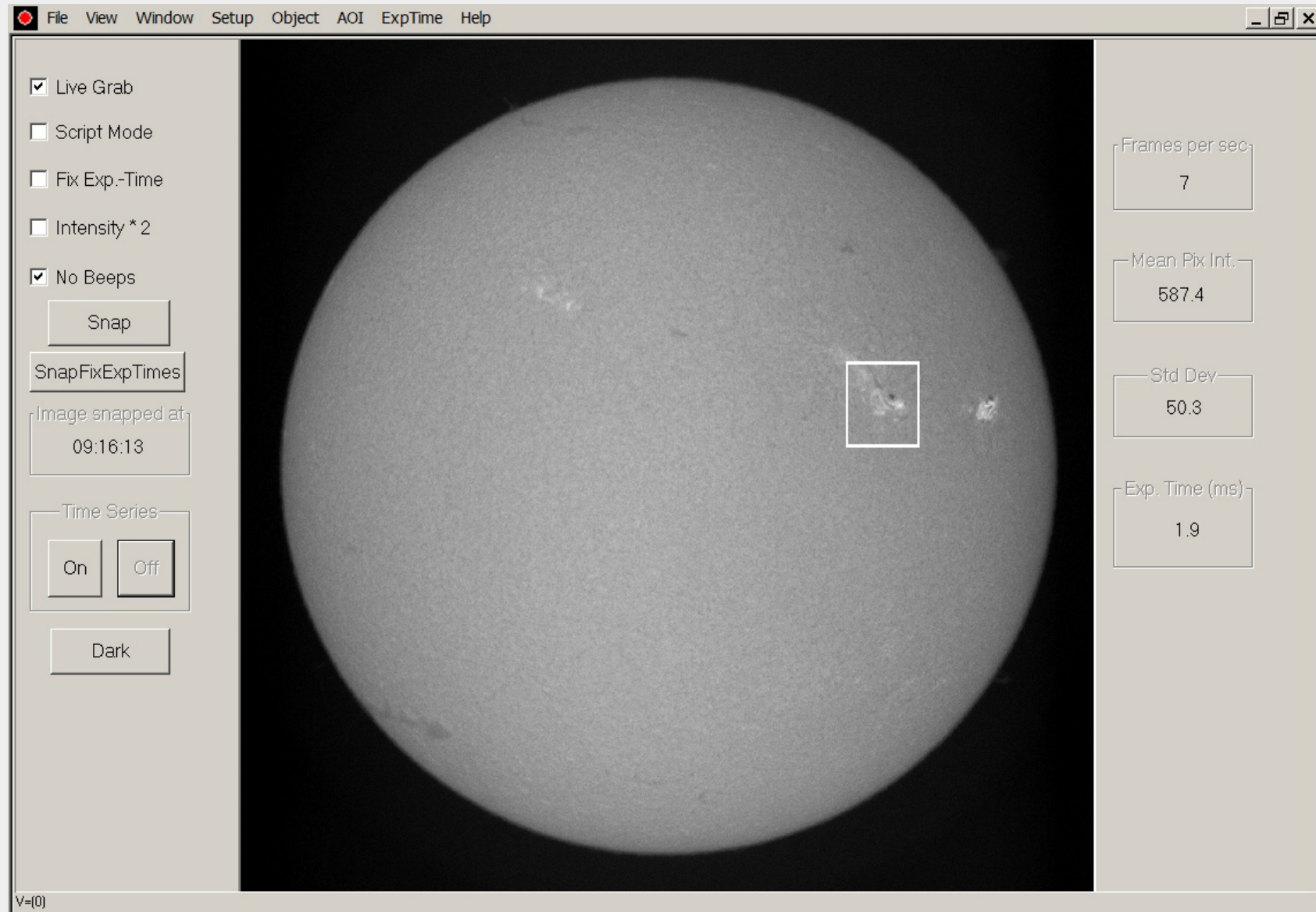
## Camera control

- each camera controlled by one PC (simple 19" server, low energy CPU, 2 x Gbit interface)
- has to withstand high temperature range (-10 to 30 deg)
- no large local storage necessary



# Camera control interface

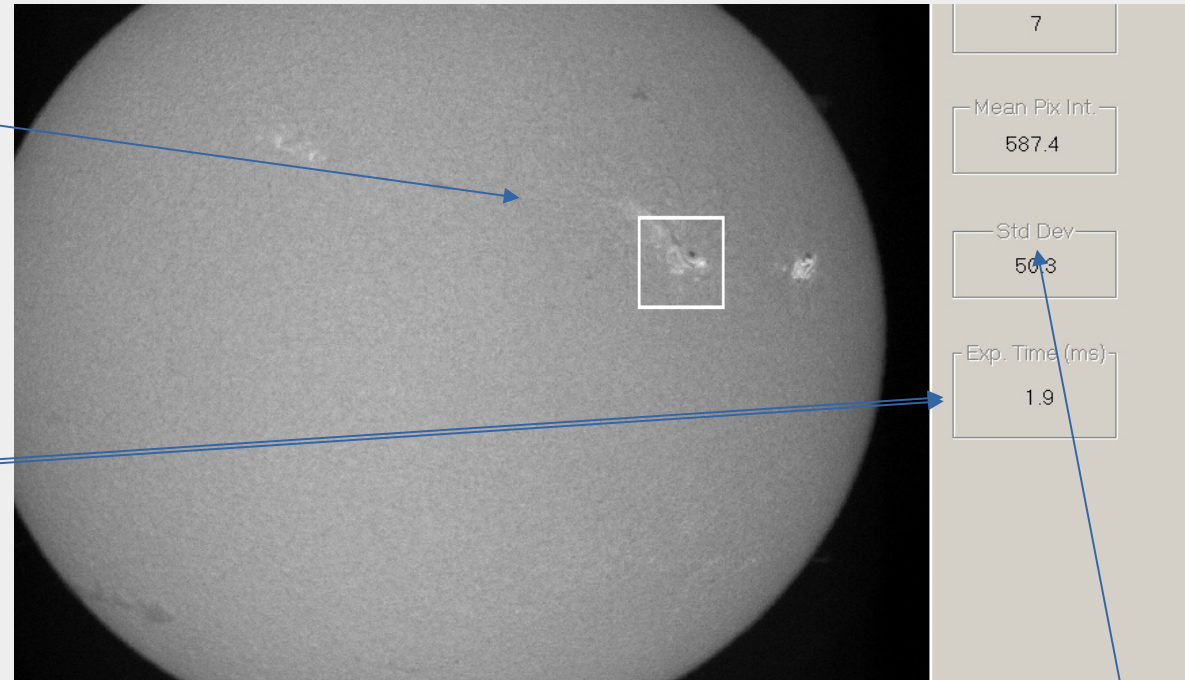
- the same for all cameras
- Visual C++ and Common Vision Blox libraries (sold with cameras)





## Camera control interface

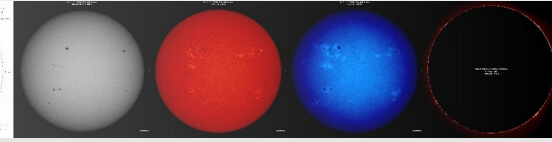
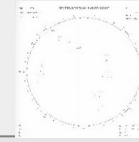
AOI- Area of Intensity  
controls exposure time and  
frame selection



**Exposure:** The mean intensity in the AOI is kept constant → flares are not overexposed  
(H-alpha = 600 counts,  
CaII K = 1200 counts,  
White-light = 2000 counts)

**Frame selection:** the image with the highest rms out of a fixed number (10) is stored → the time delay between the stored images is not always the same (5 to 7 seconds)





## acquisition – temporary archive

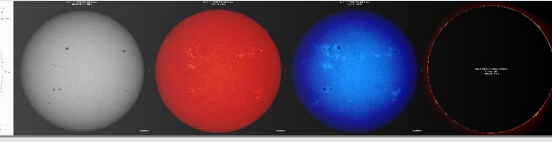
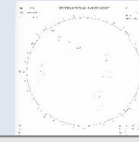
images are not stored locally

all images are stored on a RAID system in daily directories as FITS and JPEG files

image names contain type, date and time (self sorting!)

```
/caii/incoming/  
|  
|-- 20220603  
| |-- ca20220603_074912.fts  
| |-- ca20220603_074912.jpg  
| |-- ca20220603_074916.fts  
| ...  
| |  
| |-- ca20220603_105812.fts  
| |-- ca20220603_105812.jpg  
| `--dc20220603_074900.fts
```

```
/phoka/incoming/  
|  
|-- 20220603  
| |-- dc20220603_074908.FTS  
| |-- wl20220603_074912.FTS  
| |-- wl20220603_074912.jpg  
| |-- wl20220603_074916.FTS  
| ...  
| |  
| |-- wl20220603_105812.FTS  
| `-- wl20220603_105812.jpg
```



# data processing

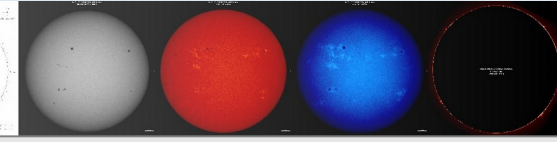
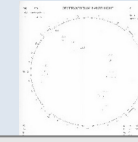
- workstation waits for new image in **incoming/** directory
- process the image
- move the processed image to **processed/** and the original to a temporary **archive/**
- the **incoming/** directory should be empty afterwards if everything runs smooth
- only images newer than the last processed image are processed (live image is always the latest grabbed image!)
- omitted images in **incoming/** (happens very seldom) are processed in the evening
- for each camera there exists a separate processing task (perl + c++)



# temporary archive

- all grabbed images are there (FITS + JPEG) also bad ones
- can hold up to one year of observations
- can be used to test quality algorithms or for machine learning
- can be used as backup if something went wrong
- „bad“ images like eclipse images can be reprocessed
- UFOs can be found!





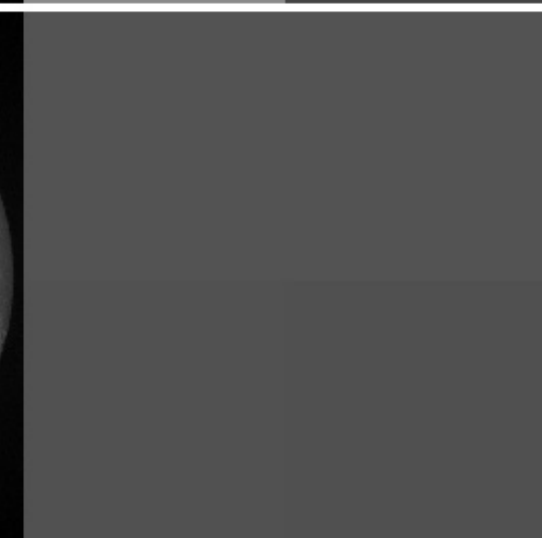
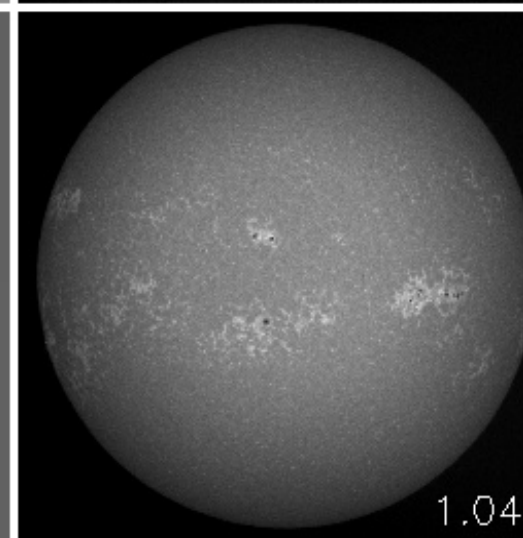
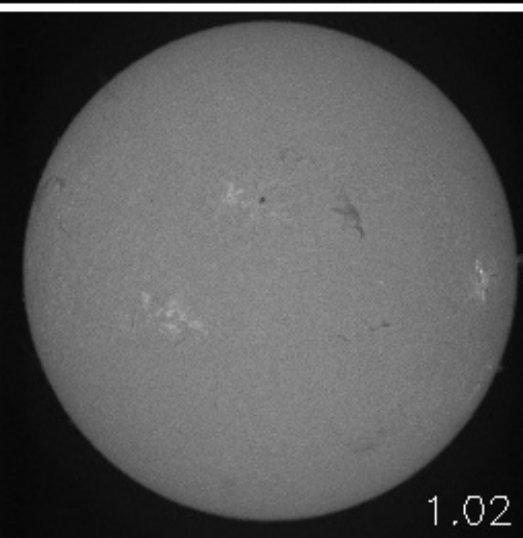
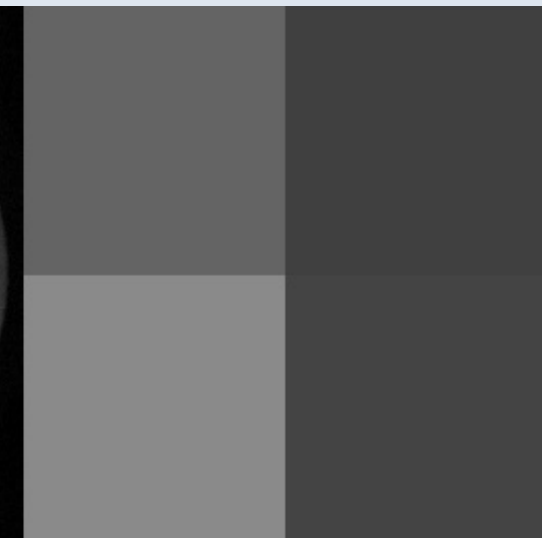
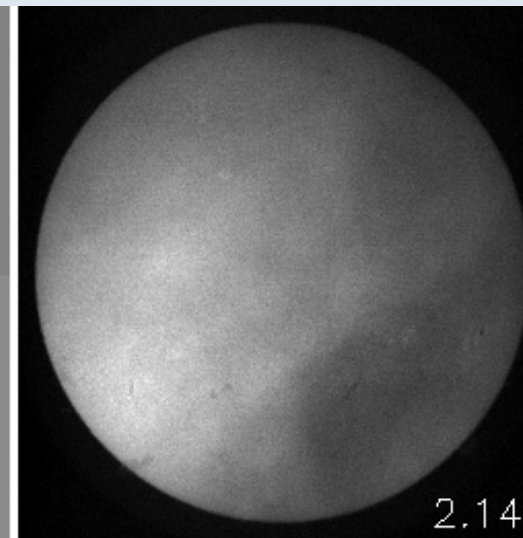
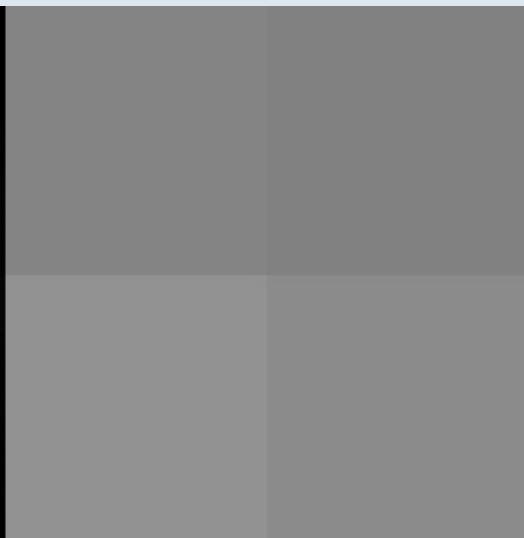
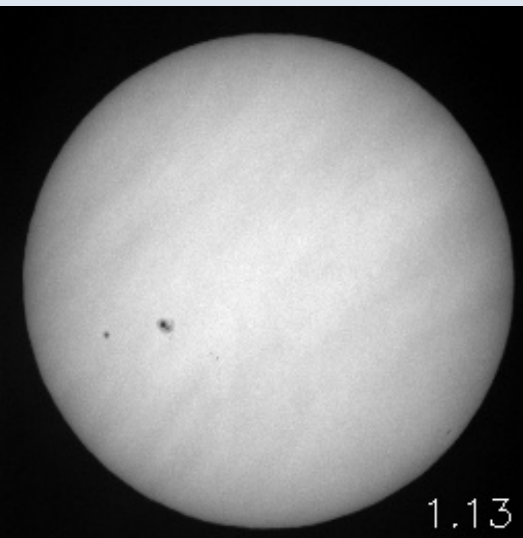
# processing – quality check

- file size (incomplete files are not used – happens with 1 out of several millions)
- exposure time: long exposure time = bad contrast (thin clouds)
- intensity in AOI – can be too high or too low when clouds move fast and camera control is slower
- inhomogeneities in image
- solar limb detection accuracy
- sharpness of image



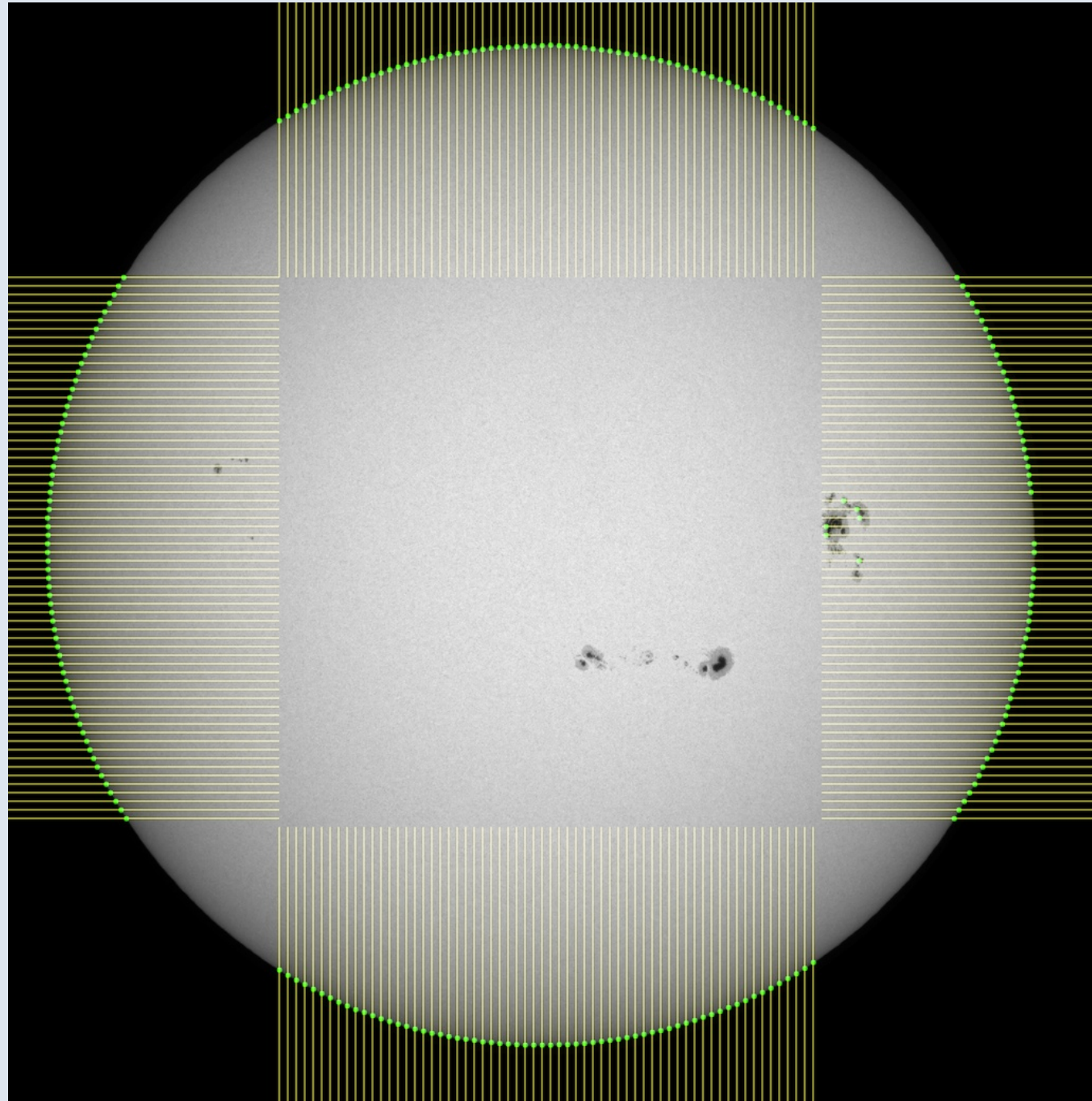
# check inhomogeneities

reduce image to 4 pixels and compare brightest to faintest



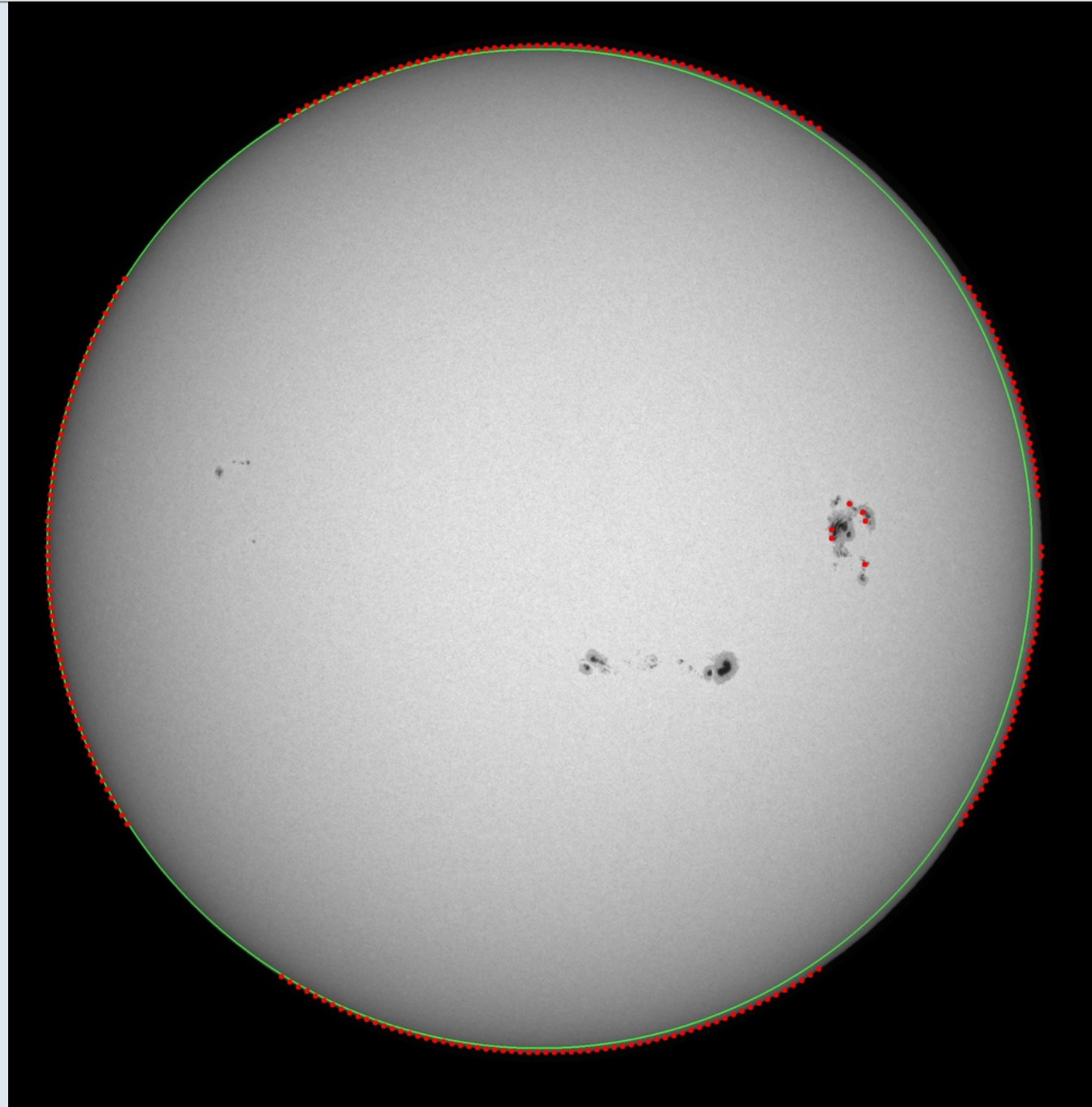
# limb detection

- find the highest intensity change in intensity profiles from the four main directions
- use these points as first guess for the solar limb and use a method like Taubin for a circle fit



# 1<sup>st</sup> iteration

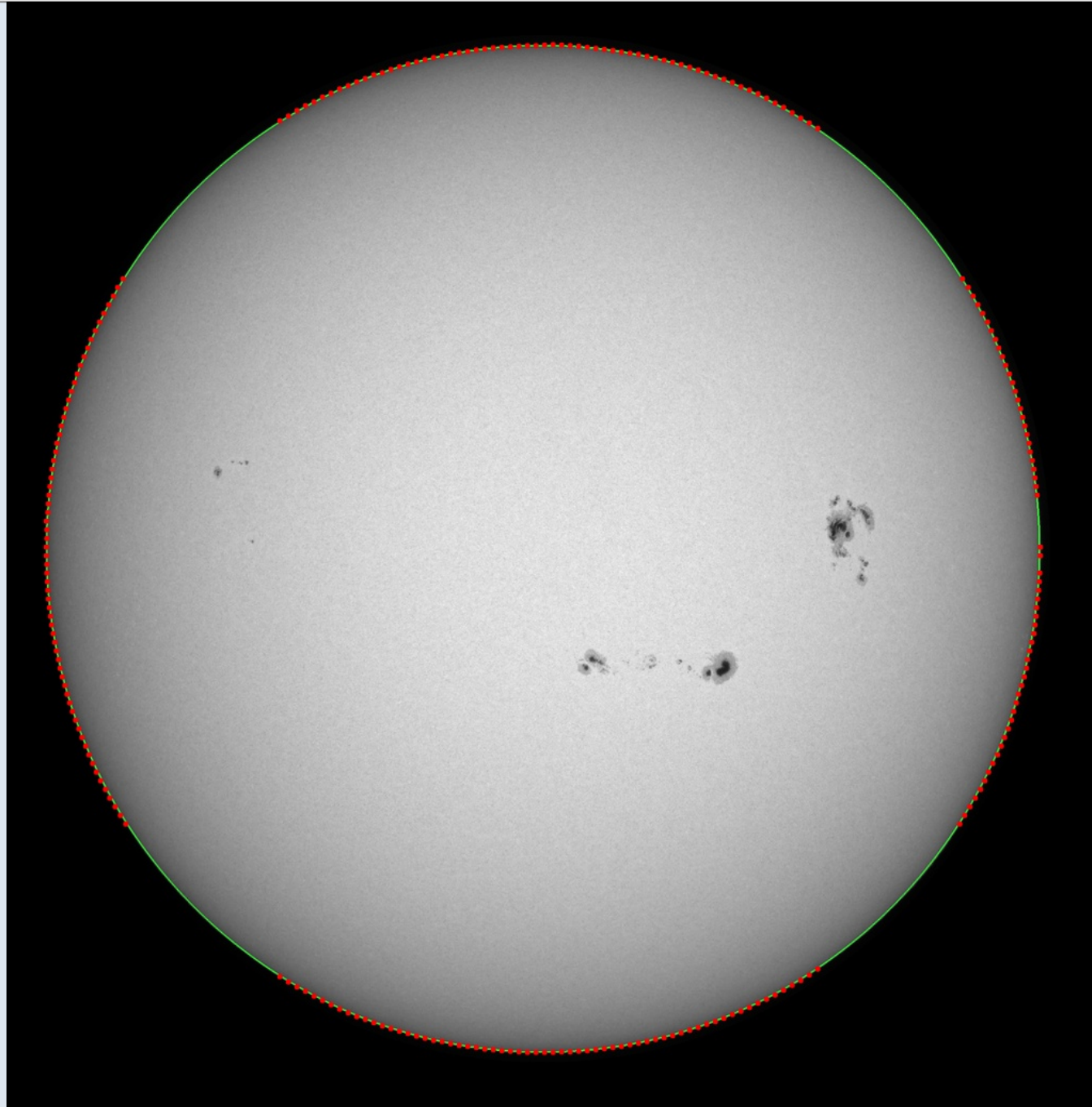
- remove all points that are further away from the mean limb (green)
- in this case the points in the sunspot are removed
- recalculate with Taubin method



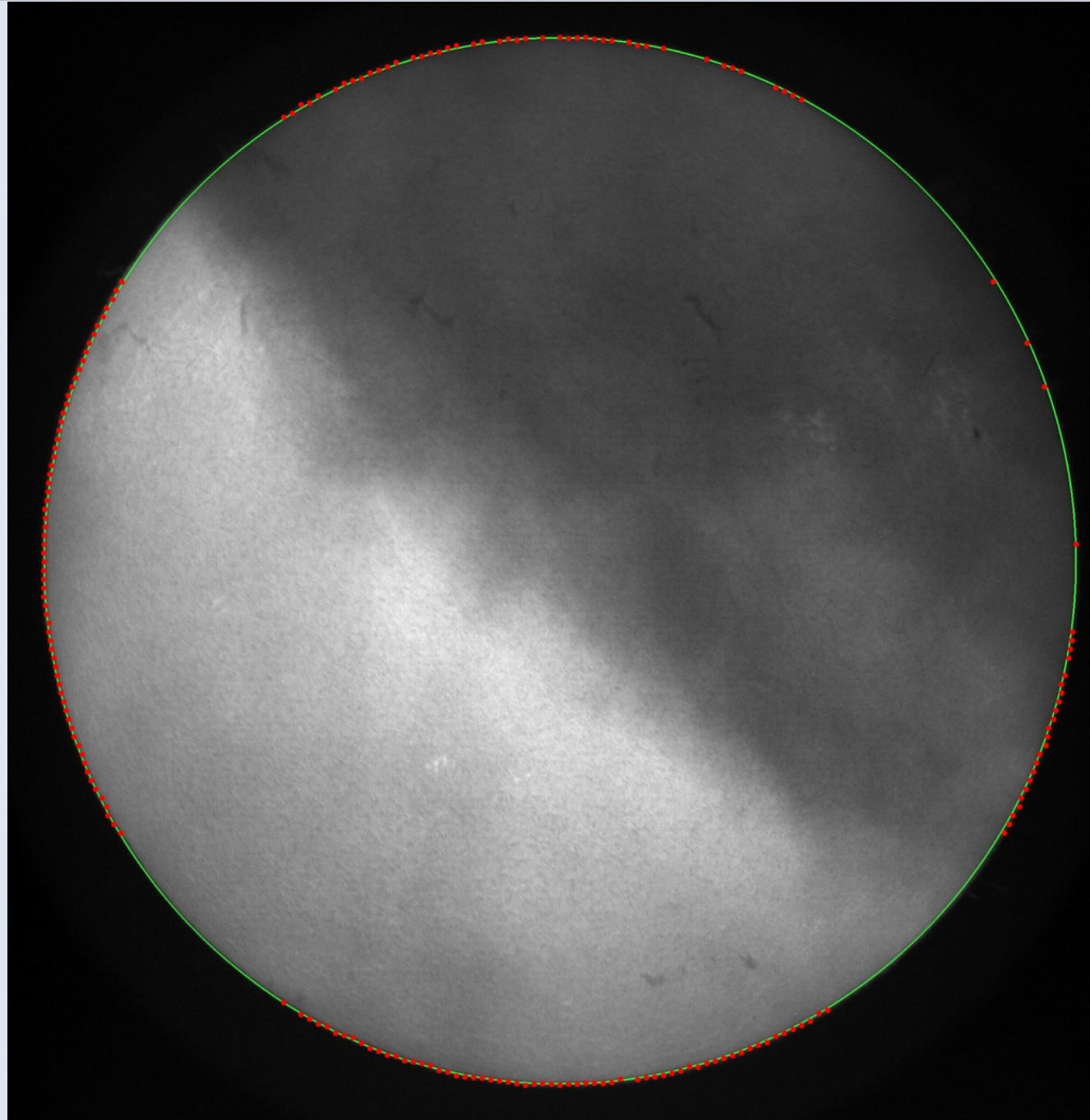


# 2<sup>nd</sup> iteration

- at least after 3 iterations the rms of the limb should be very low – if not then the image is considered to be bad!
- method is very stable and fast, small clouds are no problem

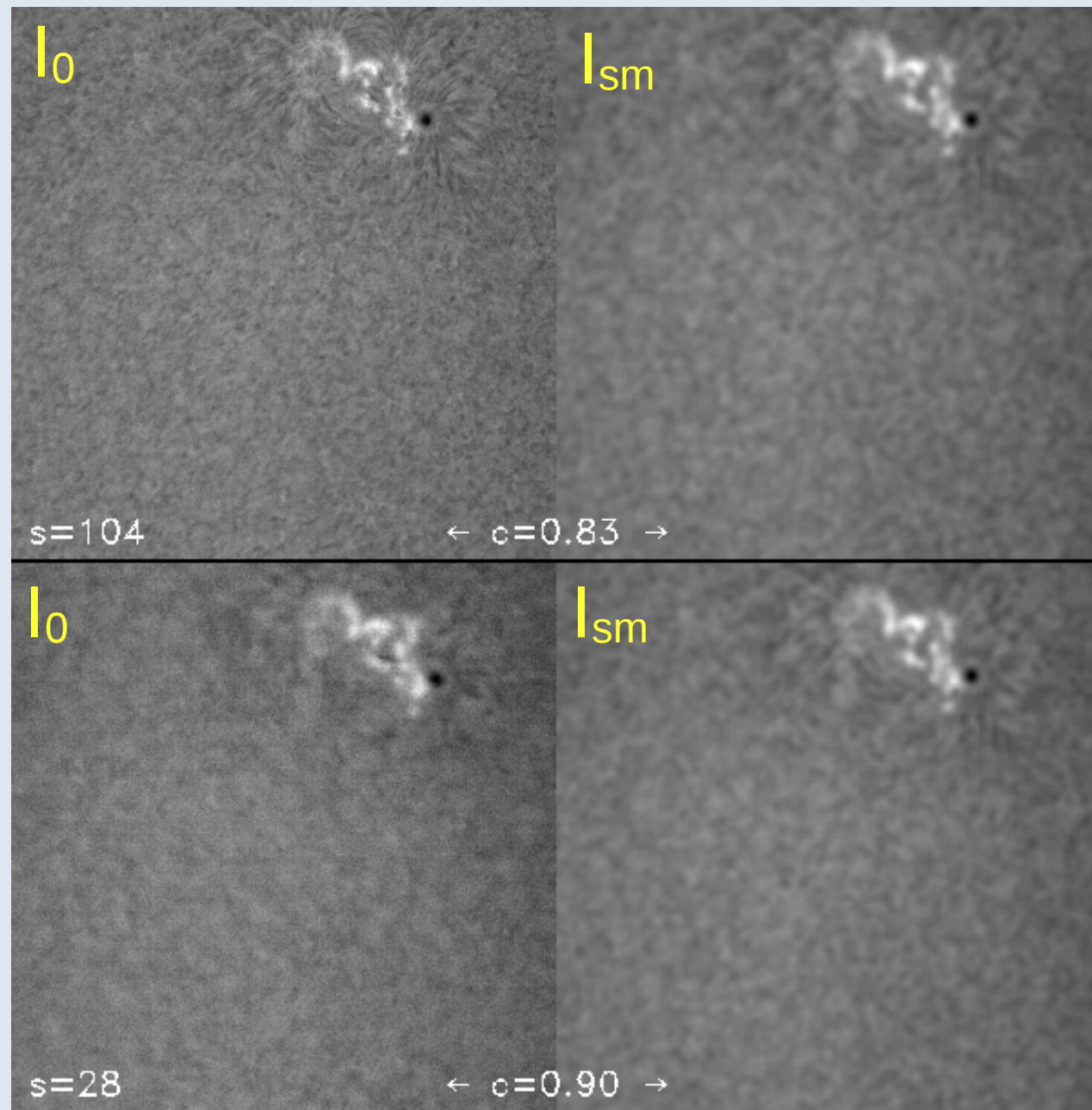


# 3 iterations with clouds!

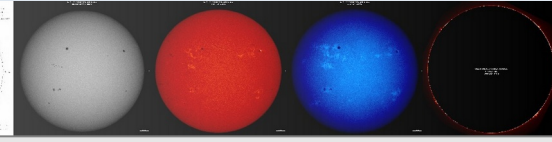
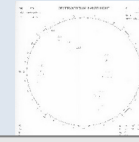


# image sharpness

- original image  $I_0$
- smooth  $I_0 \rightarrow I_{sm}$
- $c = \text{correlate}(I_0, I_{sm})$
- $c \approx 1 \rightarrow$  unsharp original image
- $c \ll 1 \rightarrow$  sharp original image







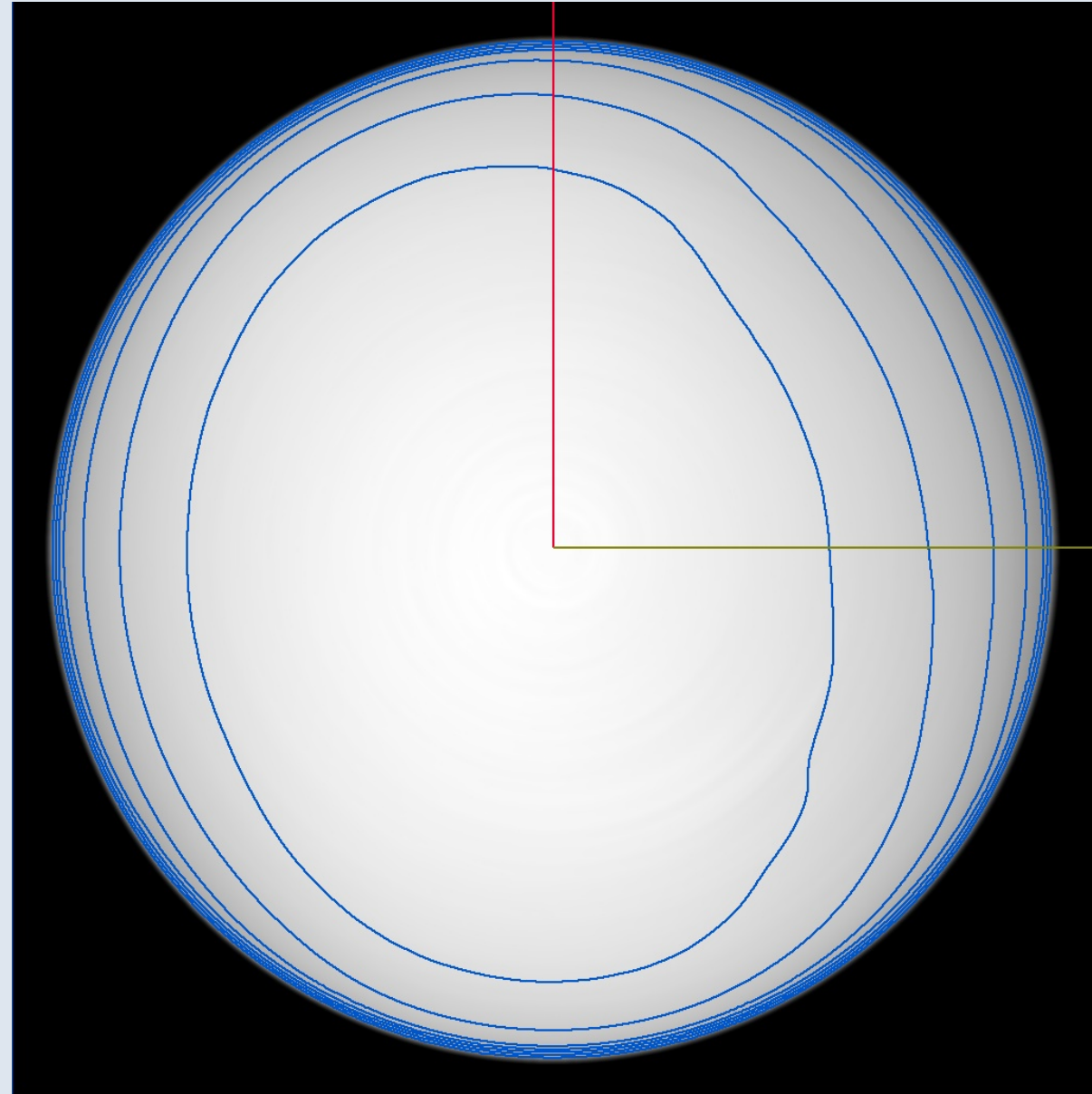
# quality decision

- combination of all values
- good = 1 → further processing, for data archive
- fair = 2 → further processed, partially stored in data archive if quality 1 is not available
- bad = 3 → no processing

# processing – CLV + anisotropic

- take concentric rings
- apply running median filter on each ring
- → map
- image/map → image without large scale anisotropies and removed CLV
- same contrast over the solar disc

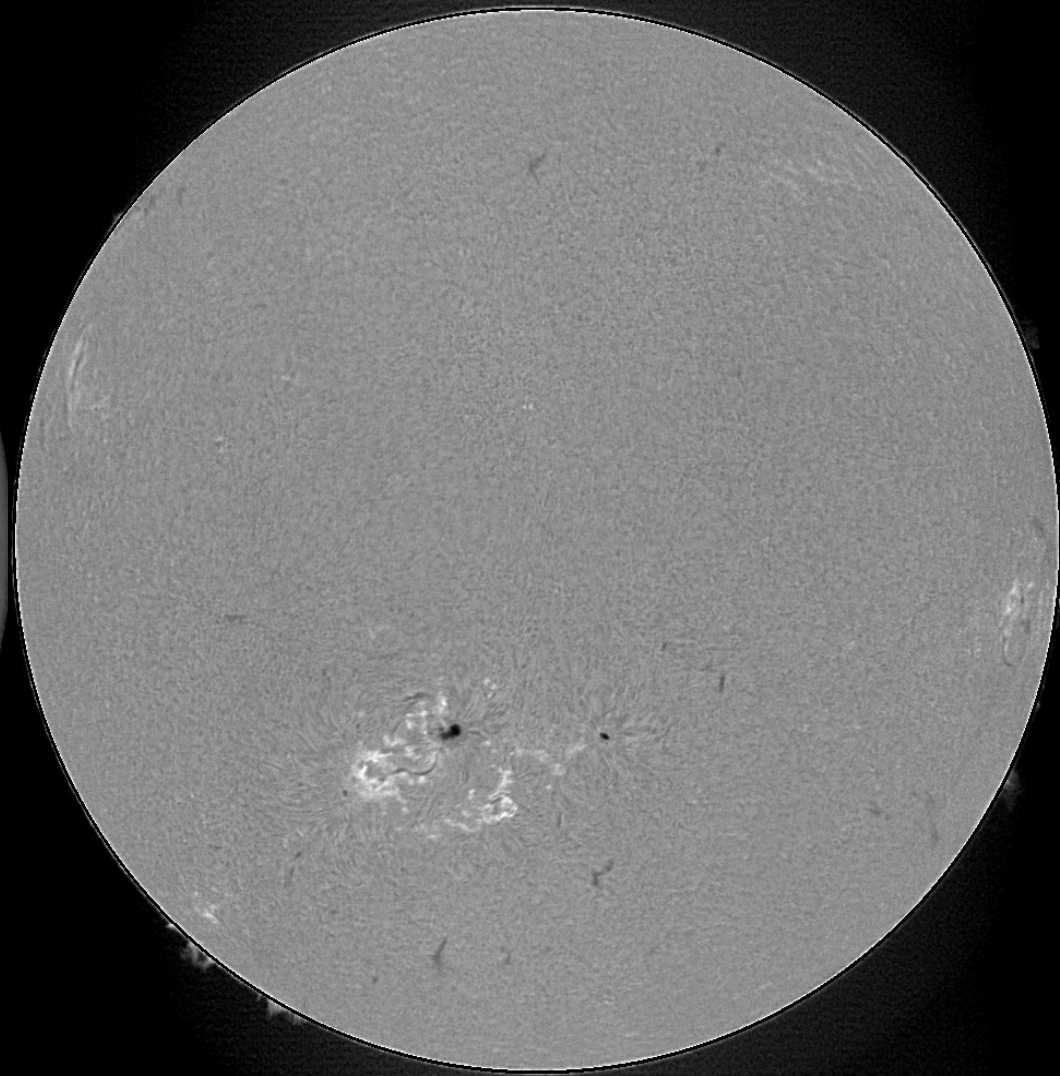
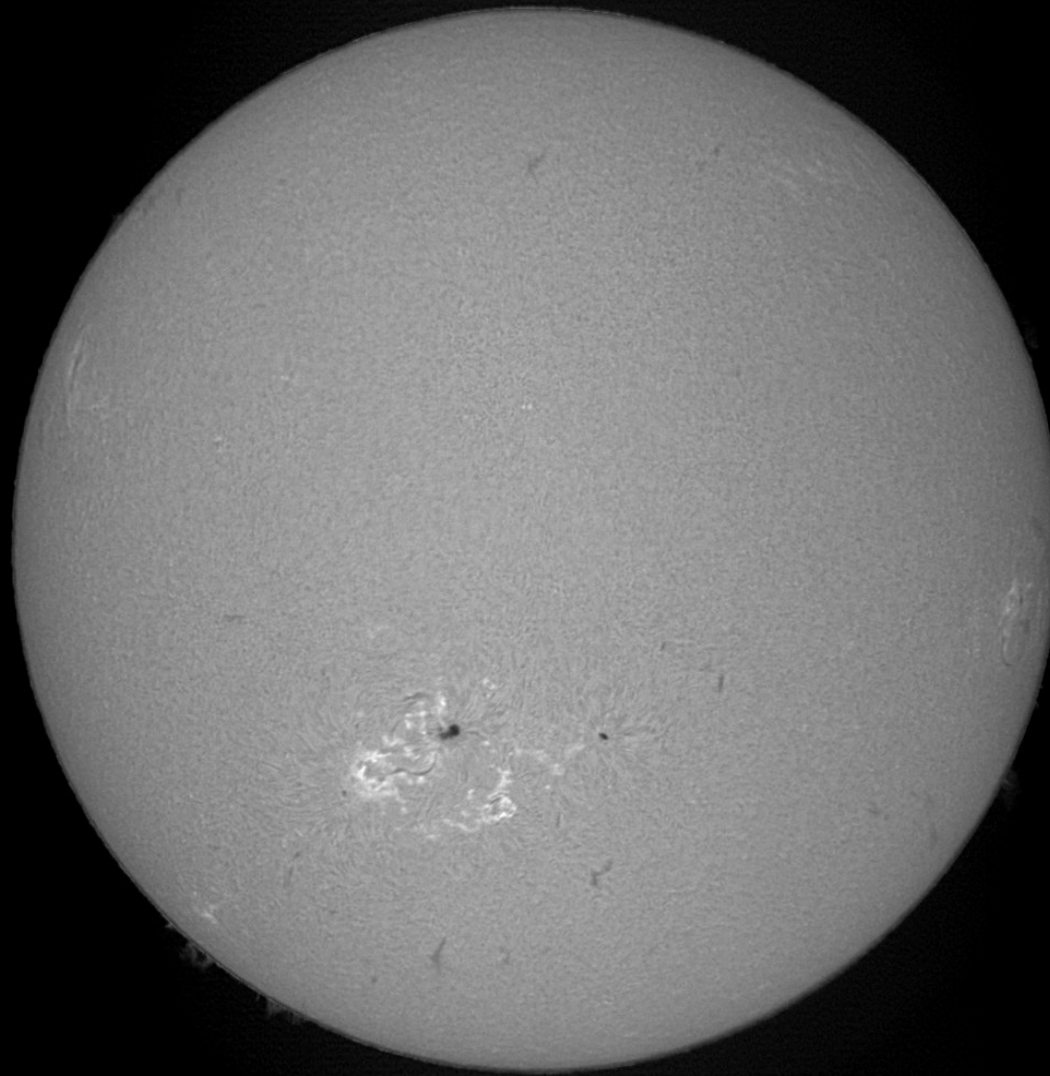
Very computation costly → parallel implementation



original

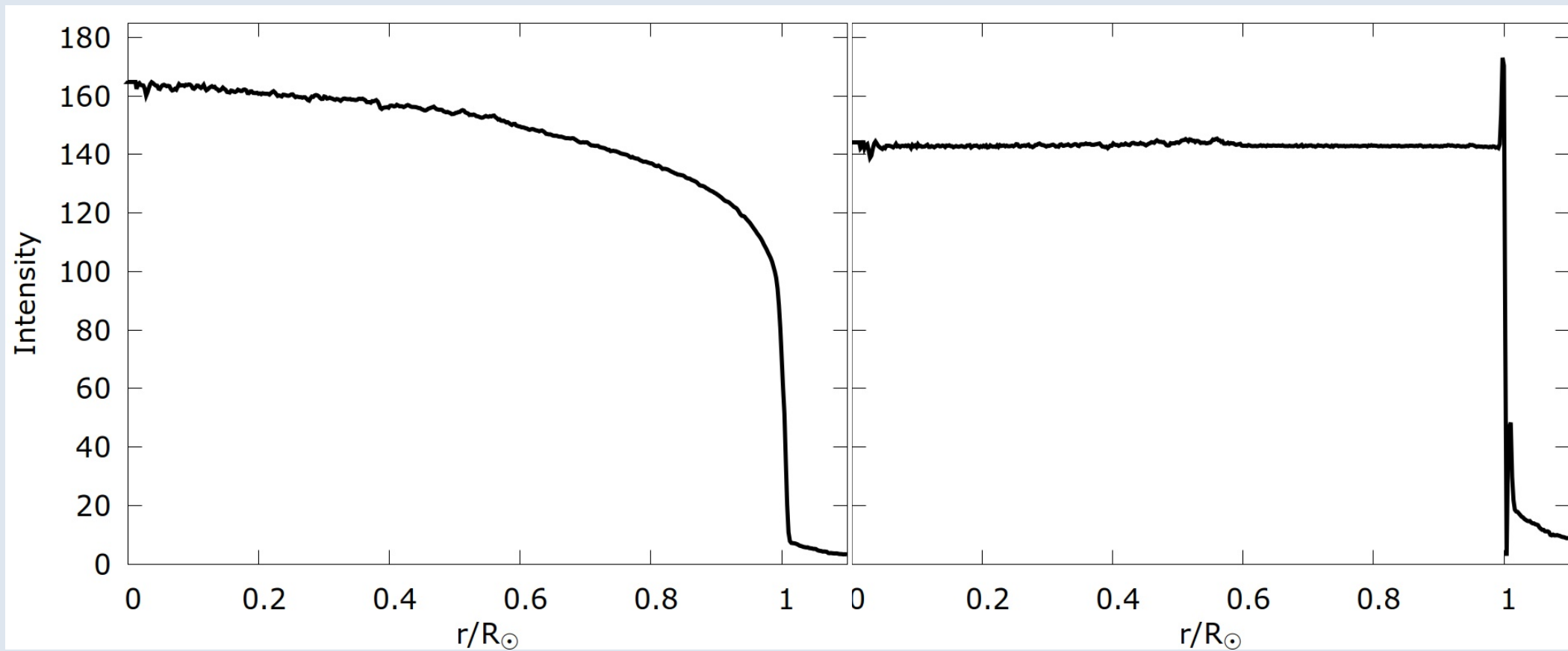


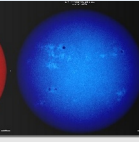
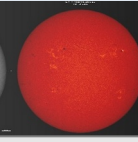
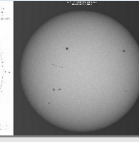
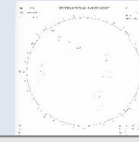
flat image





# intensity profile after correction





# file types

- temporary archive: original FITS and JPEG
- archive:
  - × processed FITS (raw data + complete header)
  - × JPEG: low contrast, high contrast, coloured with grid
  - × H-alpha: filament maps (binary and JPEG)

# data archive

Stored on 2 RAID  
systems

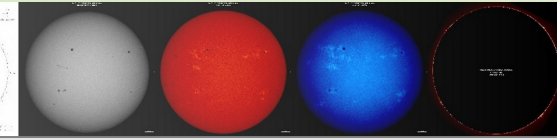
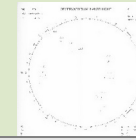
mirrored

accessible via www,  
ftp and API

no direct write  
access





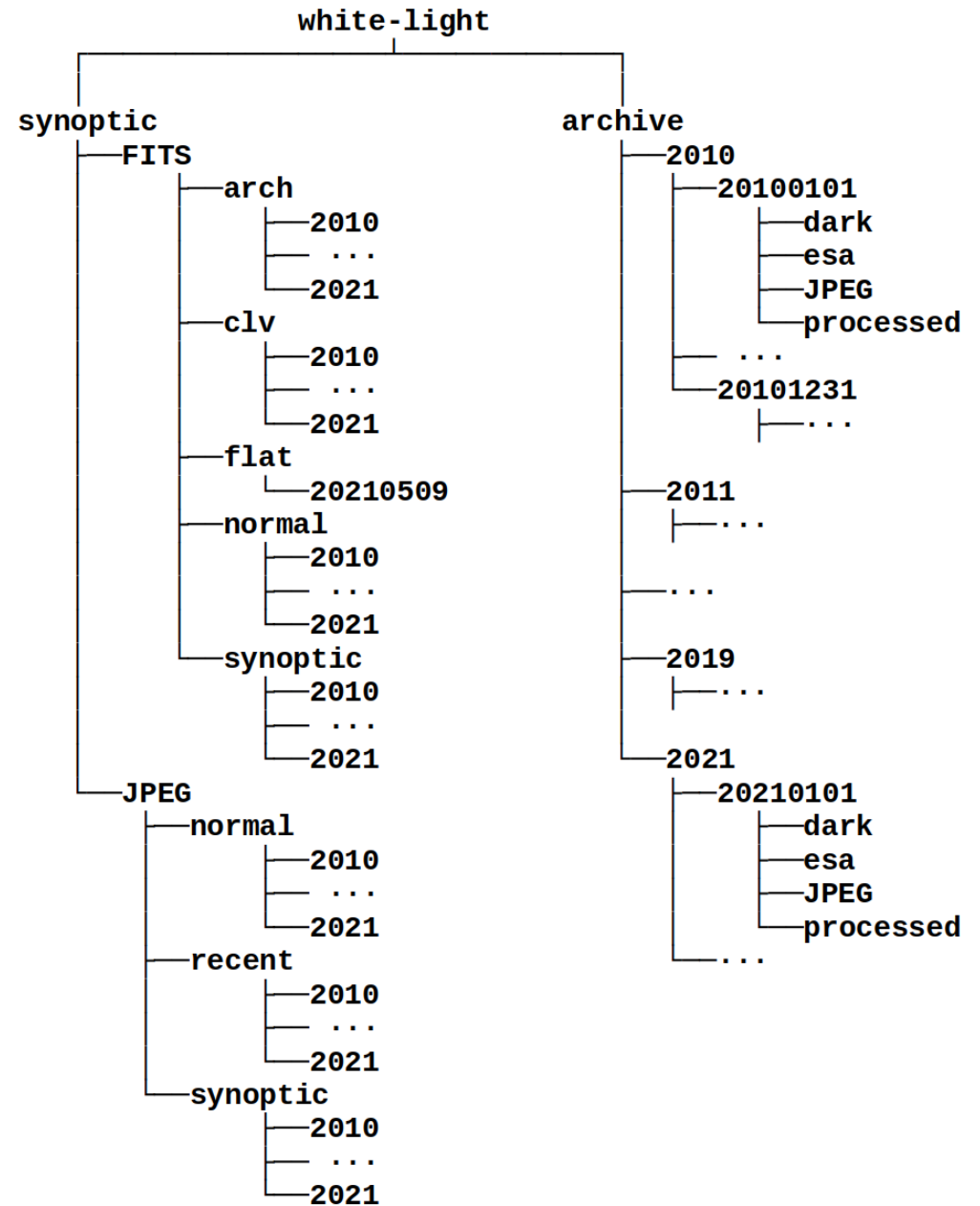


# archive structure

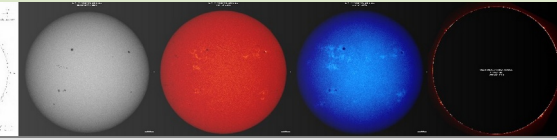
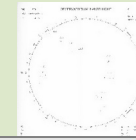
hierarchy:

1. camera
2. year
3. day

synoptic branch:  
one image per day, just  
for the www and overview







# FITS header

```

SIMPLE = T
BITPIX = 16
NAXIS = 2
NAXIS1 =2048
NAXIS2 =2048
EXTEND = 'F'
FILENAME='kanz_halph_fi_20170416_081837.fts.gz'
DATE = '2017-04-16T08:18:39'
DATE-OBS= '2017-04-16T08:18:37'
DATE-BEG= '2017-04-16T08:18:37'
TIMESYS = 'UTC'
OBSVTRY = 'Kanzelhoehe Observatory'
TELESCOP= 'KHPI'
INSTRUME= 'HA2'
DETECTOR= 'TM4200-6'
OBJECT = 'Full Sun'
FILTER = 'Zeiss Lyot Halpha'
WAVELNTH= 6563.28 / [ANG], FWHM=0.7 [ANG]
WAVEMIN = 6562.93
WAVEMAX = 6563.63
EXP_TIME= 2.752 / Exposure Time [ms]
XPOSURE = 0.002752 / [s]
BSCALE = 1
BZERO = 32768
BUNIT = 'CCD COUNTS'
DATAMIN = 0
DATAMEAN= 572
DATAMAX = 795
CTYPE1 = 'SOLAR_X'
CTYPE2 = 'SOLAR_Y'
CUNIT1 = 'arcsec'
CUNIT2 = 'arcsec'
CRPIX1 = 1024.50
CRPIX2 = 1024.50

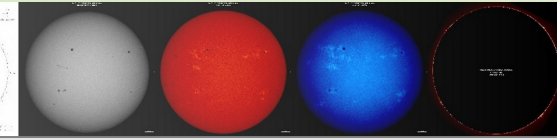
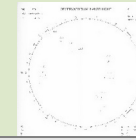
```

```

SIMPLE T for FITS Standard
BITPIX Bits per Pixel - 16bit Integer
NAXIS # ofdimensions
NAXIS1 x-dimension
NAXIS2 y-dimension
EXTEND F - normal FITS file
FILENAME ...
DATE time of processing
DATE-OBS time of image capture
DATE-BEG -"-
TIMESYS time frame
OBSVTRY
TELESCOP name of telescope
INSTRUME name of instrument
DETECTOR name of camera
OBJECT observed object
FILTER type of filter
WAVELNTH central wavelength
WAVEMIN minimum wavelength
WAVEMAX maximum wavelength
EXP_TIME exposure time in ms
XPOSURE exposure time in s
BSCALE intesity scaling
BZERO offset of intesity scaling
BUNIT units
DATAMIN minimum value
DATAMEAN mean value
DATAMAX maximum value
CTYPE1 coordinates x-axis
CTYPE2 coordinates y-axis
CUNIT1 coordinate unit x-axis
CUNIT2 coordinate unit y-axis
CRPIX1 center coordinate x
CRPIX2 center coordinate y

```





# FITS header

```

CDEL1 = 1.019699 / [arcsec/pix]
CDEL2 = 1.019699 / [arcsec/pix]
CRVAL1 = 0.9184746
CRVAL2 = -16.49616
ANGLE = -26.68136 / [deg]
CROT1 = 26.68136 / [deg]
CENTER_X = 1023.599 / [pix]
CENTER_Y = 1040.677 / [pix]
SOLAR_R = 942.486 / [pix]
RSUN_REF = 6.9938E+08 / [m]
RSUN_ARC = 960.859 / [arcsec]
SOLAR_P0 = -25.92443 / [deg]
SOLAR_B0 = -5.503757 / [deg]
CAR_ROT = 2189
QUALITY = 2 / image quality [1-3]
OBS_TYPE = 'HALPH '
OBS_PROG = 'HALPHA PATROL'
TYPE-DP = 'ARCHIVE ' / Data Processing Type
EXP_MODE = 0 / Exp. Mode
           (0=auto,1=dbl,2=fix,3=both)
PRE_INT = 600 / Preselected PixInt in AOI
A_O_INT = '1223,714,1223,1032' / Rect. for PixInt [X0,Y0,X1,Y1]
ORIGIN = 'KANZELHOEHE OBSERVATORY, A-9521 TREFFEN, AUSTRIA
COMMENT Orientation: N up, W right, first pix is left bottom
HISTORY No intensity processing applied
END

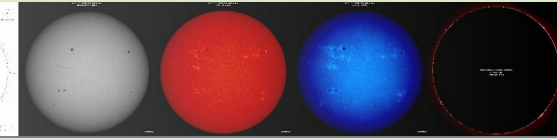
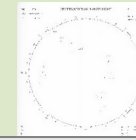
```

```

CDEL1 pixel/arcsec x-coordinates
CDEL2 pixel/arcsec y-coordinates
CRVAL1 center coordinate Sun x
CRVAL2 center coordinate Sun y
ANGLE rotation of solar N (cameratilt+P0)
CROT1 -rotation of solar N (cameratilt+P0)
CENTER_X Sun center x
CENTER_Y Sun center y
SOLAR_R solar radius in pixel
RSUN_REF solar radius in m
RSUN_ARC solar radius in arcsec
SOLAR_P0 solar north
SOLAR_B0 solar ecliptic tilt
CAR-ROT Carrington rotation number
QUALITY quality
OBS_TYPE observation type
OBS_PROG observation programme
TYPE-DP data type
EXP_MODE exposure mode

PRE_INT fixed intensity in AOI
A_O_INT AOI coordinates
ORIGIN data source
COMMENT
HISTORY
END

```



# Database - mysql

phpMyAdmin

Letzte Favoriten

- Neu
- gaestebuch
- information\_schema
- KEAS
  - Neu
  - caii
  - detector
  - dha\_dd\_corr
  - file\_type
  - filter
  - flares
  - hafilm
  - hafilter\_steps\_wavelen
  - halpha
  - instrument
  - instrument\_detector
  - ISN
  - kfactor
  - korona
  - logbook
  - object\_list
  - observation\_type
  - observer
  - phoka
  - scientist
  - seasons
  - seeing\_quality
  - ssd
  - sunspots
  - syn\_data
  - techlog
  - wl
- library

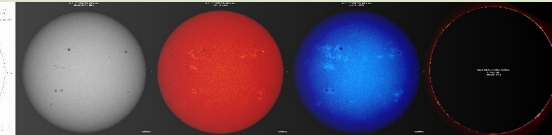
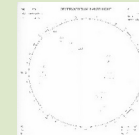
Server: localhost > Datenbank: KEAS

Struktur SQL Suche Abfrage Exportieren Importieren Operationen Rechte Routinen Ereignisse Trigger

Filter

Beinhalten das Wort:

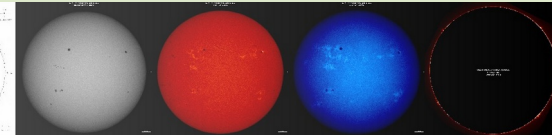
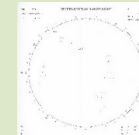
Table	Aktion	Datensätze	Typ	Kollation	Größe	Überhang
<input type="checkbox"/> caii	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	~1.099.001	InnoDB	utf8_general_ci	185,7 MiB	-
<input type="checkbox"/> detector	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	16	InnoDB	utf8_general_ci	16,0 KiB	-
<input type="checkbox"/> dha_dd_corr	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	366	InnoDB	utf8_general_ci	16,0 KiB	-
<input type="checkbox"/> file_type	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	11	InnoDB	utf8_general_ci	16,0 KiB	-
<input type="checkbox"/> filter	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	37	InnoDB	utf8_general_ci	16,0 KiB	-
<input type="checkbox"/> flares	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	17.887	InnoDB	utf8_general_ci	1,5 MiB	-
<input type="checkbox"/> hafilm	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	~465.103	InnoDB	utf8_general_ci	70,6 MiB	-
<input type="checkbox"/> hafilter_steps_wavelen	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	33	InnoDB	utf8_general_ci	16,0 KiB	-
<input type="checkbox"/> halpha	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	~3.847.540	InnoDB	utf8_general_ci	775,0 MiB	-
<input type="checkbox"/> instrument	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	12	InnoDB	utf8_general_ci	16,0 KiB	-
<input type="checkbox"/> instrument_detector	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	17	InnoDB	utf8_general_ci	16,0 KiB	-
<input type="checkbox"/> ISN	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	~75.686	InnoDB	utf8_general_ci	2,5 MiB	-
<input type="checkbox"/> kfactor	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	5	InnoDB	utf8_general_ci	16,0 KiB	-
<input type="checkbox"/> korona	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	4.139	InnoDB	utf8_general_ci	144,0 KiB	-
<input type="checkbox"/> logbook	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	~68.234	InnoDB	utf8_general_ci	5,5 MiB	-
<input type="checkbox"/> object_list	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	2	InnoDB	utf8_general_ci	16,0 KiB	-
<input type="checkbox"/> observation_type	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	14	InnoDB	utf8_general_ci	16,0 KiB	-
<input type="checkbox"/> observer	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	22	InnoDB	utf8_general_ci	16,0 KiB	-
<input type="checkbox"/> phoka	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	13.417	InnoDB	utf8_general_ci	1,5 MiB	-
<input type="checkbox"/> scientist	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	57	InnoDB	utf8_general_ci	16,0 KiB	-
<input type="checkbox"/> seasons	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	124	InnoDB	utf8_general_ci	16,0 KiB	-
<input type="checkbox"/> seeing_quality	★ <a href="#">Anzeigen</a> <a href="#">Struktur</a> <a href="#">Suche</a> <a href="#">Einfügen</a> <a href="#">Leeren</a> <a href="#">Löschen</a>	9	InnoDB	utf8_general_ci	16,0 KiB	-



# H-alpha data

exposure_key	detector_key	filter_key	object_list_key	start_date	exposure_time	quality	xcen	ycen	radius	xsize	ysize	filesize	filename	esa
7407384	16	20	2	2023-10-23 13:57:02	4.443	1	1024.349	1023.244	947.440	2048	2048	4927528	halpha3/archive/2023/20231023/processed/kanz_halph...	3
7407981	16	20	2	2023-10-23 13:56:50	4.831	1	1020.908	1024.154	946.674	2048	2048	4932136	halpha3/archive/2023/20231023/processed/kanz_halph...	3
7407980	16	20	2	2023-10-23 13:56:38	4.680	1	1021.847	1023.014	947.236	2048	2048	4930194	halpha3/archive/2023/20231023/processed/kanz_halph...	3
7407979	16	20	2	2023-10-23 13:56:13	3.639	2	1022.583	1024.794	947.043	2048	2048	4933620	halpha3/archive/2023/20231023/processed/kanz_halph...	0
7407383	16	20	2	2023-10-23 13:56:01	3.492	1	1022.304	1023.731	947.656	2048	2048	4964729	halpha3/archive/2023/20231023/processed/kanz_halph...	3
7407978	16	20	2	2023-10-23 13:55:55	3.756	1	1022.963	1023.343	946.908	2048	2048	4999160	halpha3/archive/2023/20231023/processed/kanz_halph...	3
7407977	16	20	2	2023-10-23 13:55:37	3.956	1	1024.125	1024.850	946.760	2048	2048	4933366	halpha3/archive/2023/20231023/processed/kanz_halph...	3
7407976	16	20	2	2023-10-23 13:55:12	4.815	1	1026.333	1024.996	947.370	2048	2048	4927278	halpha3/archive/2023/20231023/processed/kanz_halph...	3
7407382	16	20	2	2023-10-23 13:55:00	5.200	1	1024.552	1024.417	947.252	2048	2048	4943398	halpha3/archive/2023/20231023/processed/kanz_halph...	3
7407975	16	20	2	2023-10-23 13:54:54	5.332	1	1024.696	1024.692	947.278	2048	2048	4967745	halpha3/archive/2023/20231023/processed/kanz_halph...	3
7407381	16	20	2	2023-10-23 13:54:30	8.080	1	1024.318	1025.541	947.586	2048	2048	4954357	halpha3/archive/2023/20231023/processed/kanz_halph...	3
7407974	16	20	2	2023-10-23 13:53:35	5.966	1	1024.013	1023.895	947.512	2048	2048	4959240	halpha3/archive/2023/20231023/processed/kanz_halph...	3
7407973	16	20	2	2023-10-23 13:53:16	6.613	2	1022.615	1025.615	947.294	2048	2048	5043637	halpha3/archive/2023/20231023/processed/kanz_halph...	0
7407380	16	20	2	2023-10-23 13:53:05	9.771	1	1022.605	1024.709	947.509	2048	2048	4975489	halpha3/archive/2023/20231023/processed/kanz_halph...	3
7407972	16	20	2	2023-10-23 13:52:52	12.656	2	1023.331	1024.408	947.422	2048	2048	4991216	halpha3/archive/2023/20231023/processed/kanz_halph...	0
7407971	16	20	2	2023-10-23 13:52:33	13.503	2	1029.143	1024.484	947.204	2048	2048	4933993	halpha3/archive/2023/20231023/processed/kanz_halph...	0
7407970	16	20	2	2023-10-23 13:52:16	12.493	2	1022.488	1025.388	947.505	2048	2048	4945384	halpha3/archive/2023/20231023/processed/kanz_halph...	0
7407379	16	20	2	2023-10-23 13:51:32	5.712	1	1025.029	1023.492	947.566	2048	2048	4940544	halpha3/archive/2023/20231023/processed/kanz_halph...	3





# Data access – cesar.kso.ac.at



## Kanzelhöhe Observatory Synoptic Archives

KARL-FRANZENS-UNIVERSITÄT GRAZ  
UNIVERSITY OF GRAZ



Navigation

Data

Docs & Info

About

Links

2023-10-24

Julian Date: 2460240.0001

$P = 25.729^\circ$   $B_0 = 5.318^\circ$   $L_0 = 76.64^\circ$  Carr# 2276

12:00 UTC

Dist. = 0.995 AU App. Diam. = 1929.0" Elevation = 30.10°

Sunrise 05:31:17 UTC Sunset 16:05:49 UTC

» Data Policy

### Live Data

- » Latest Images
- » Latest H $\alpha$  (ESA SSA)
- » Flare Monitoring (live)

### Synoptic Data

- » Daily Overview
- » Two Weeks Photosphere
- » Two Weeks Chromosphere
- » Sunspot Numbers
- » Sunspot Drawings
- » H $\alpha$
- » White-light
- » CaIIK
- » Filaments

### Full Data Archive

- » Archive / Ftp-Server / Local
- » Fast Mirror Archive / Graz

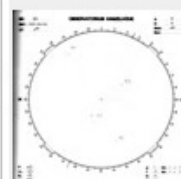
### Archive Database Search

- » Observation Database
- » Sunspot Numbers
- » Flares (visually)
- » Flares (automatically)
- » Flares (NOAA-SWPC)
- » KSO observing logs

### Misc.

- » Solar Ephemeris
- » Debrecen Photohel. Data

### Sun - Photosphere for 2023-10-22



#### Kanzelhöhe Sunspot Drawing

06:53 UTC

Side reversed!

» Drawing Archive



#### Kanzelhöhe White-light

06:52:23 UTC

Synoptic \* .jpg \* .fits.gz

Contrast \* .jpg \* .fits.gz

Movie

» Synoptic Archive

» Full White-light Archive

$R_1 = 52$

$g = 4$

$f = 12$

Observing Log

Monthly Summary

NOAA/SWPC Event List

H $\alpha$ /GOES Intensity Plot

### Sun - Chromosphere for 2023-10-22



#### Kanzelhöhe H $\alpha$

06:53:31 UTC

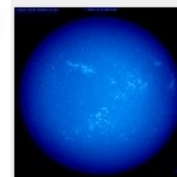
Synoptic \* .jpg \* .fits.gz

Contrast \* .jpg \* .fits.gz

Movie

» Synoptic Archive

» Full H $\alpha$  Archive



#### Kanzelhöhe Ca II K

06:52:15 UTC

Synoptic \* .jpg \* .fits.gz

Contrast \* .jpg \* .fits.gz

Movie

» Synoptic Archive

» Full Ca II K Archive



#### Kanzelhöhe H $\alpha$ Prominence Images

06:56:46 UTC

» Synoptic Archive

« 2023-10-21

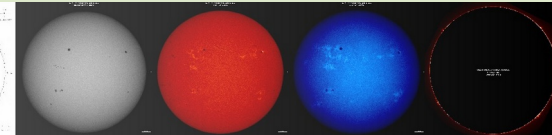
« 1 Month

2023-10-22

2023-10-23 »

### Panorama: 2023-10-22 12:00 (UTC)





# Data access – [kanzelhohe.uni-graz.at](http://kanzelhohe.uni-graz.at)






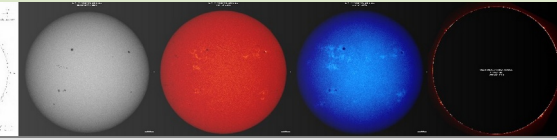
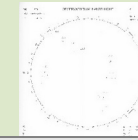
Observatorium Kanzelhöhe  
für Sonnen- und Umweltforschung

KARL-FRANZENS-UNIVERSITÄT GRAZ  
UNIVERSITY OF GRAZ



## Kanzelhöhe H $\alpha$ , Whitelight and CaIIK Data Archive

H $\alpha$ Archive	Whitelight Archive	CaIIK Archive
 <p data-bbox="234 1077 606 1101">Bad weather conditions - dome closed</p> <p data-bbox="244 1412 595 1436">(latest H<math>\alpha</math> image from KSO)</p>	 <p data-bbox="872 1077 1244 1101">Bad weather conditions - dome closed</p> <p data-bbox="851 1412 1276 1436">(latest Whitelight image from KSO)</p>	 <p data-bbox="1521 1077 1893 1101">Bad weather conditions - dome closed</p> <p data-bbox="1510 1412 1904 1436">(latest CaIIK image from KSO)</p>
<a href="#">» Archive...</a>	<a href="#">» Archive...</a>	<a href="#">» Archive...</a>



# Data access – ftp.kso.ac.at

```
/userhomes/poetzi> ftp ftp.kso.ac.at
Connected to halley.kso.ac.at.
220 Welcome on Kanzelhoehe ftp server
Name (ftp.kso.ac.at:poetzi): download
331 Password required for download
Password:
230-Welcome to the Kanzelhoehe Observatory FTP server!
This account allows just to download files, don't try anything else!
In case of problems contact poetzi@kso.ac.at!
```

CaII	calcium archive data	2010 -
caii	calcium synoptic data	2010 -
HaFilm	Ha scanned films 1024x1024 @8bit	1973 - 2000
Ha1Mold	Ha 1 MPixel archive 1008x1016 @8bit	1998 - 2005
Ha1M	Ha 1 MPixel archive 1000x1012 @10bit	2005 - 2010
halpha	Ha 1 Mpixel synoptic (8bit and 10bit)	1997 - 2010
Ha4M	Ha 4 MPixel archive 2024x2048 @12bit	2008 -
halpha2k	Ha BBSO 4 MPixel synoptic	2000 -
halpha4M	Ha 4 Mpixel synoptic	2008 -
korona	Korona Images of Ha 4 Mpixel	2009 -
Phoka	Whitelight 4 MPixel archive 2024x2048	2007 -
phokaD	Whitelight 4 Mpixel synoptic	2007 -
phokascan	Scanned Whitelight images	1989 - 2007
KSO_flares	Flare data from KSO	1984 -
Patrol_Times	Patrol time data (graph and text)	2000 -
sunspots	Drawings, Sunspot Numbers	1944 -
uvb	UV data archive	2004 - 2013

```
230 User download logged in
Remote system type is UNIX.
Using binary mode to transfer files.
ftp> █
```